

# Project SPIRAL2-CZ in NPI CAS

Jaromír Mrazek



EUROPEAN UNION  
European Structural and Investment Funds  
Operational Programme Research,  
Development and Education



# What is GANIL/SPIRAL2

1975 - Decision to build „GANIL“

1983 – first experiment

**GANIL –**  
**Grande Accélérateur**  
**National**  
**d'Ion Lourdes**



# GANIL

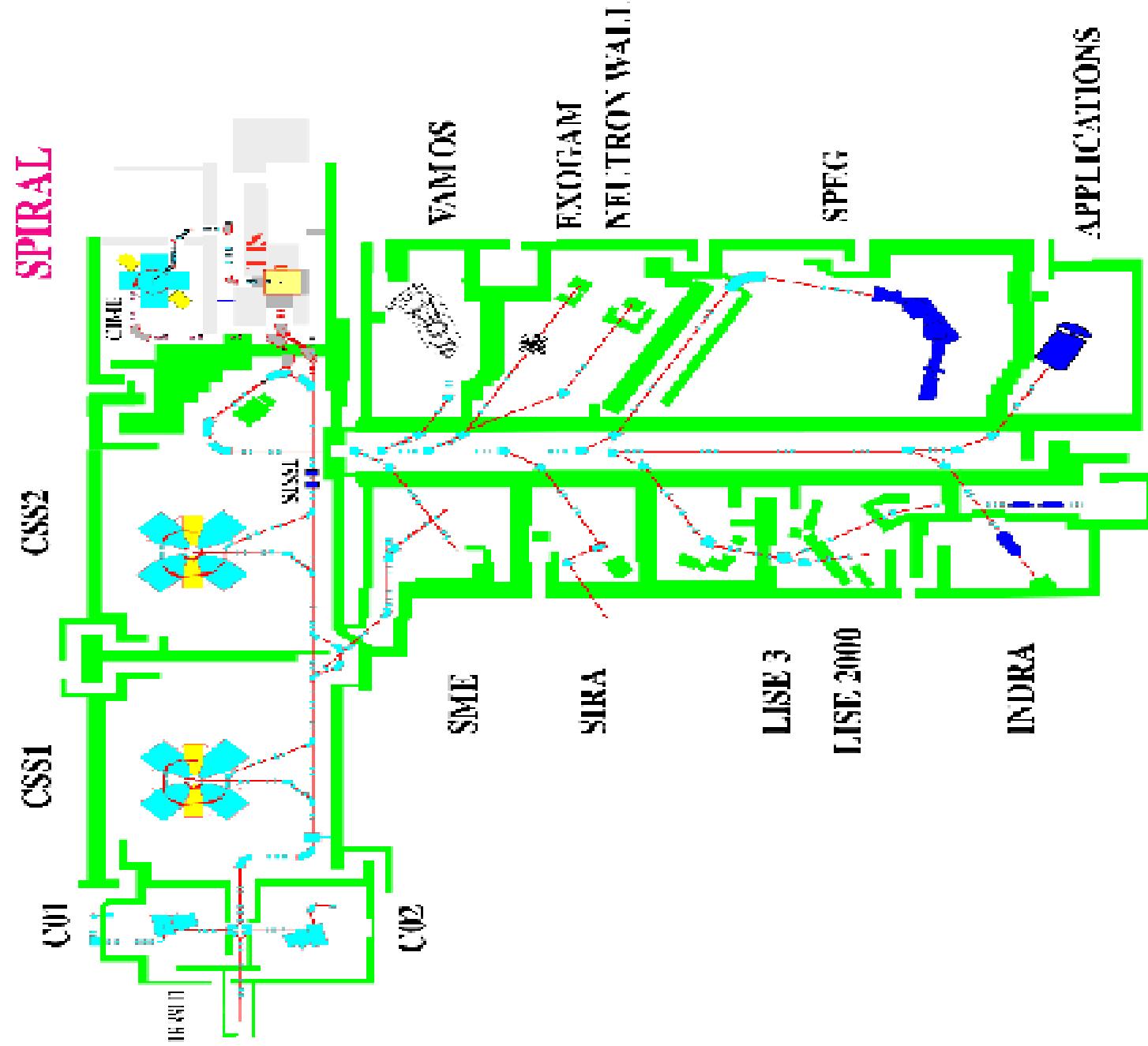
Two cyclotrons CSS1 CSS2

- stable beams

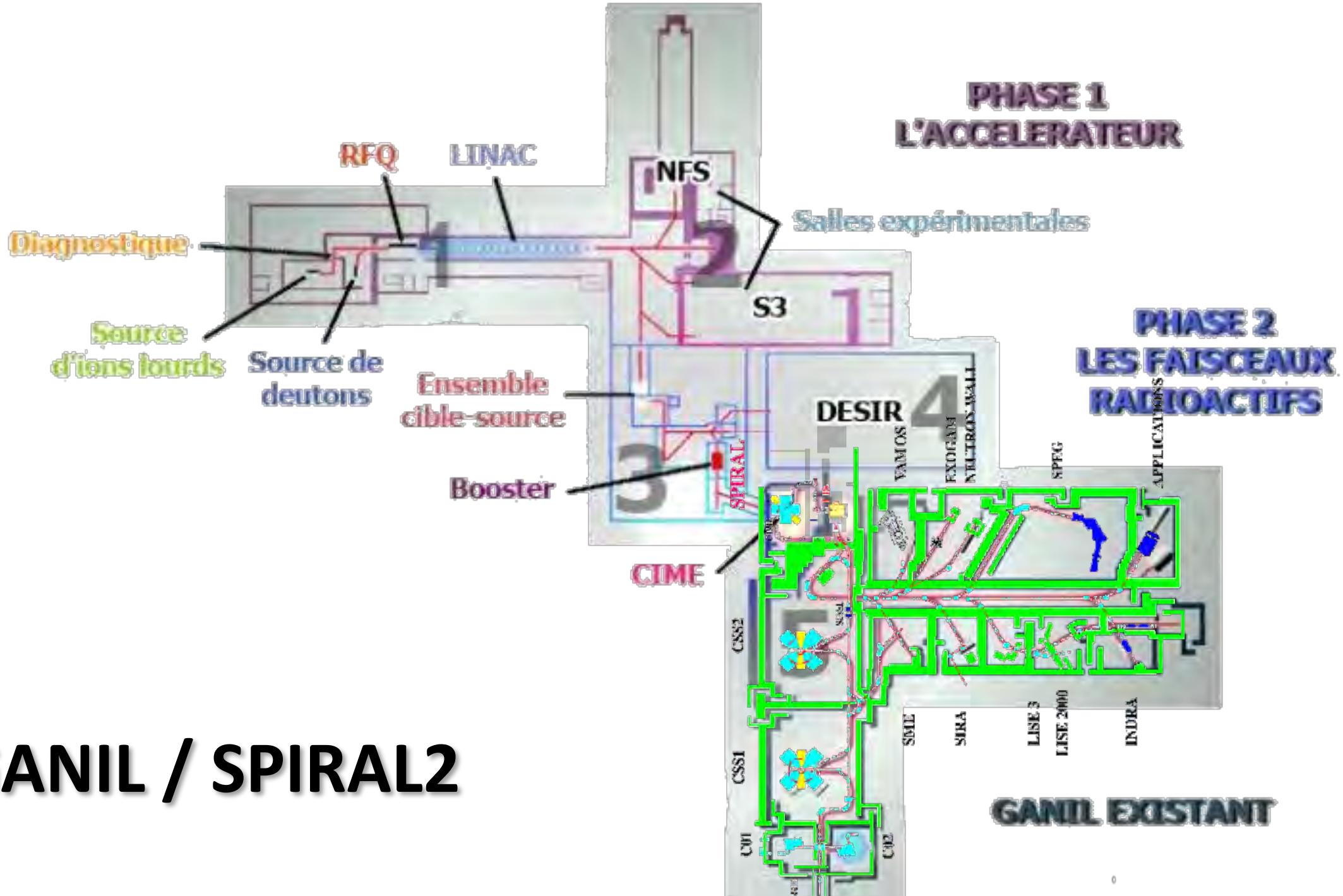
- beams by fragmentation

One post-acceleration cyclotron  
for radioactive species  
at low energies

“SPIRAL”



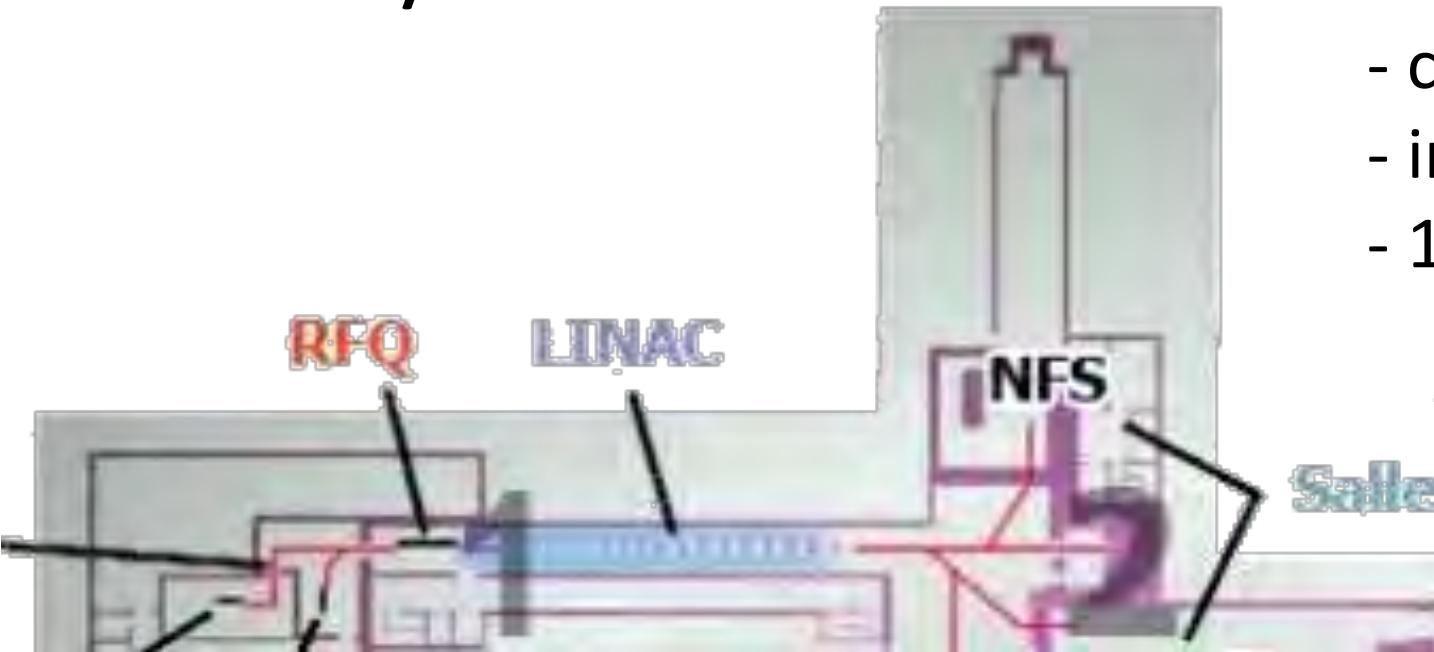
# GANIL / SPIRAL2



# SPIRAL2/NFS presented by X.Ledoux

**Primary goal – neutron beams**

**2018/2019**



Since

- close to LINAC
- intensive beams (limit 50uA / 5mA)
- 1st equipment to be commissioned

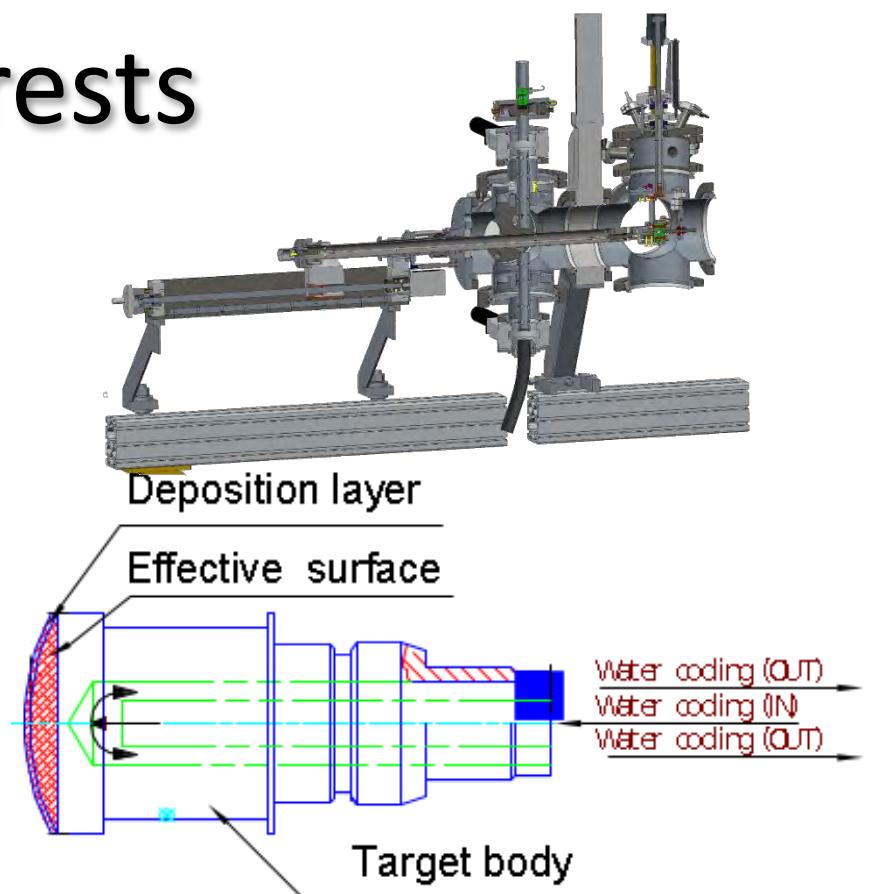
**Secondary goal**

- charged particle experiments

# Czech Republic interests

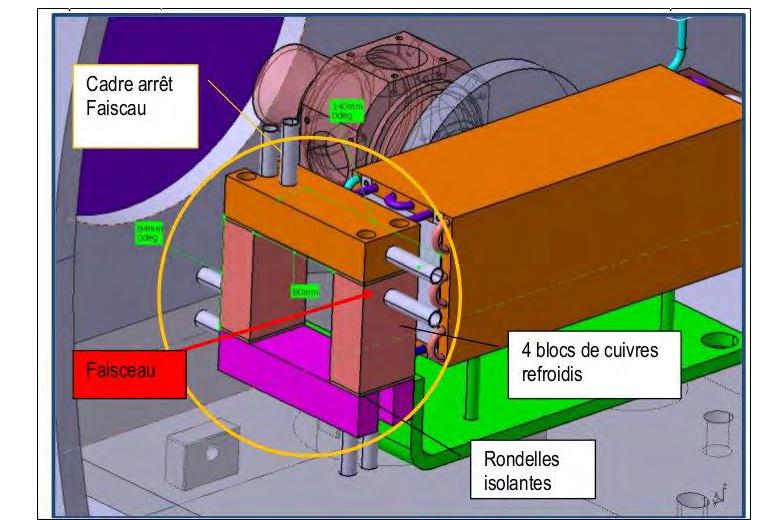
## Program of Activation by charged particles

- E.Simeckova, NPI CAS, M+V Avrigeanu, IFIN HH



## Radioisotopes for medicine

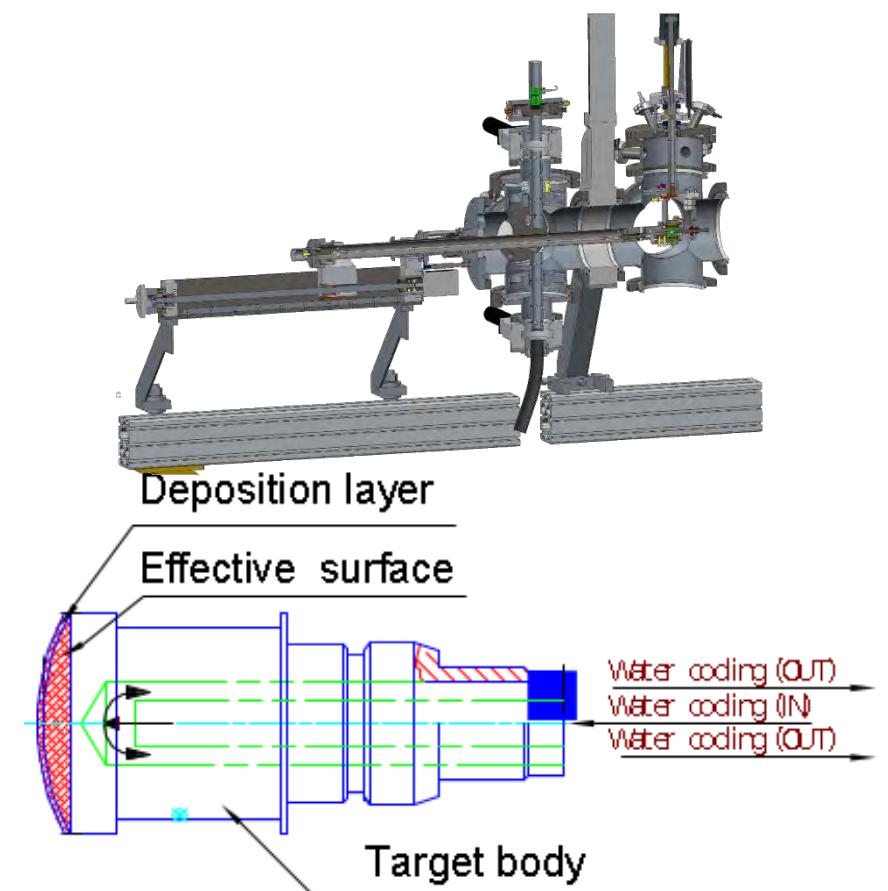
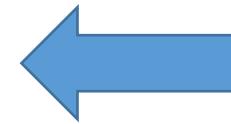
- O.Lebeda, G. de France



## SPIRAL2-CZ

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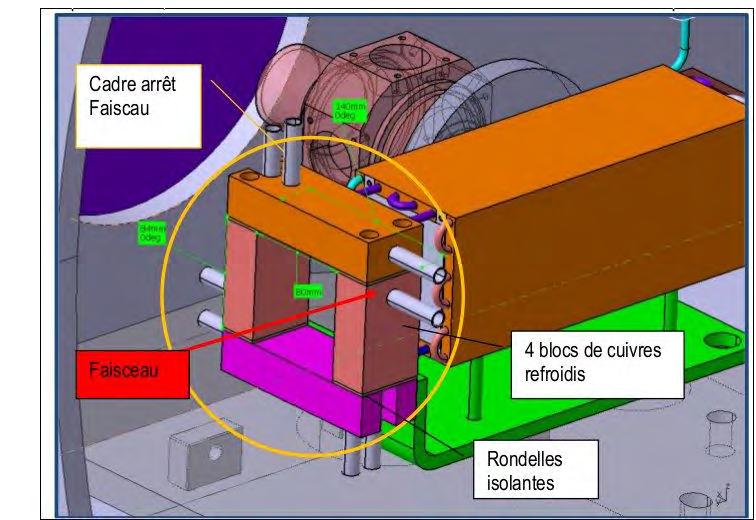
- O.Lebeda, G. de France

## Nuclear astrophysics

- J.Mrazek, F.de Oliveira, B.Bastin



# SPIRAL2-CZ



## SPIRAL2-CZ

has appeared on  
the roadmap of large research infrastructures of Czech Republic 2016-2022

Plan of support from MEYS (plan 7 years)

LM SPIRAL2-CZ (4 y., to 7 y)  
- infrastructure  
development

SPIRAL2-CZ-OP (3.r)  
Operational program EU  
- Investments  
- research



EUROPEAN UNION  
European Structural and Investment Funds  
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## Activation by charged particles in NFS

- experimental equipment: irradiation chamber`+ PTS

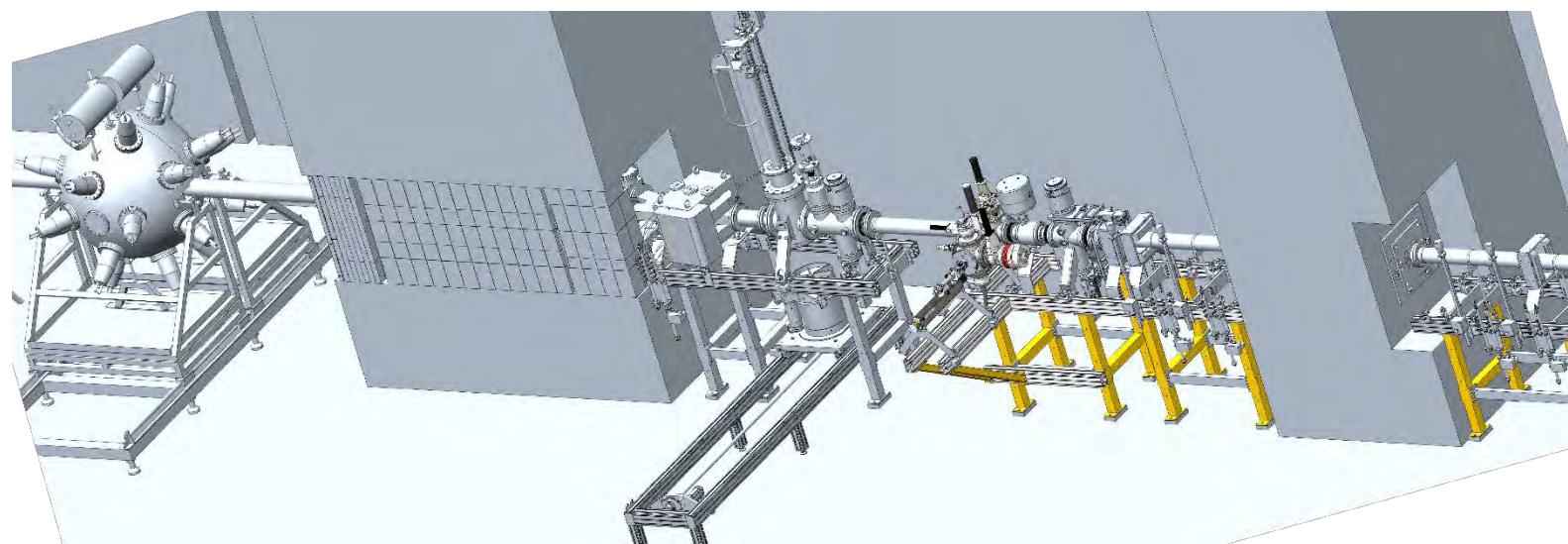
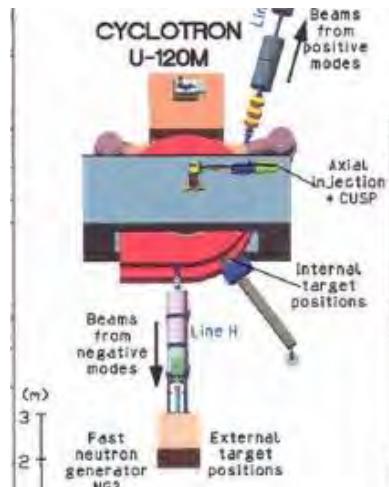
## Experiments at low energies in NFS

- nuclear astrophysics context

# Activation by charged particles: motivation

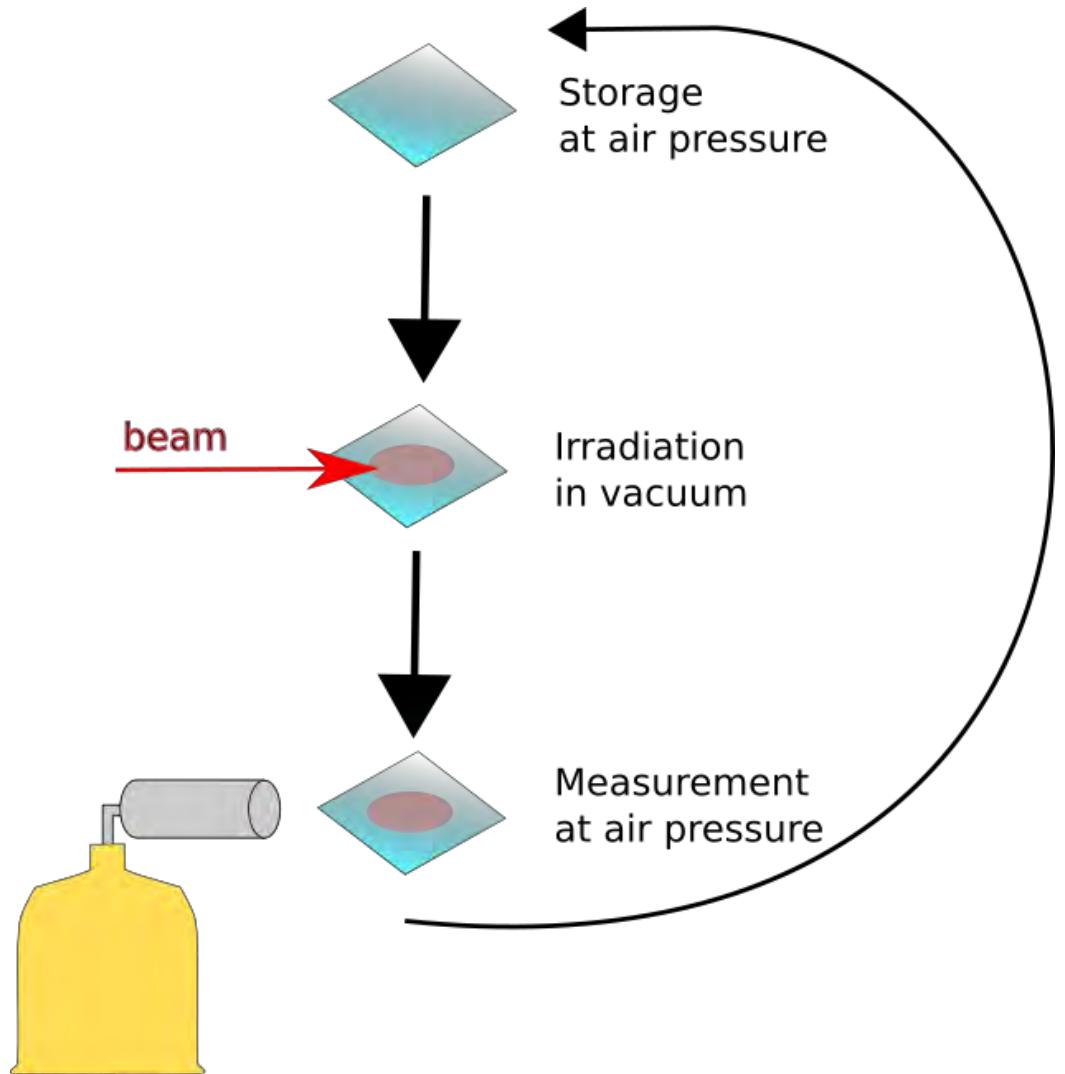
presentation of M.Avrigeanu

- **NPI CAS Rez**
- deuterons in up to 20 MeV
- half-lives limited by stack handling
- **SPIRAL2/NFS**
- deuterons up to 40 MeV
- short-lived isotopes



# Activation by charged particles

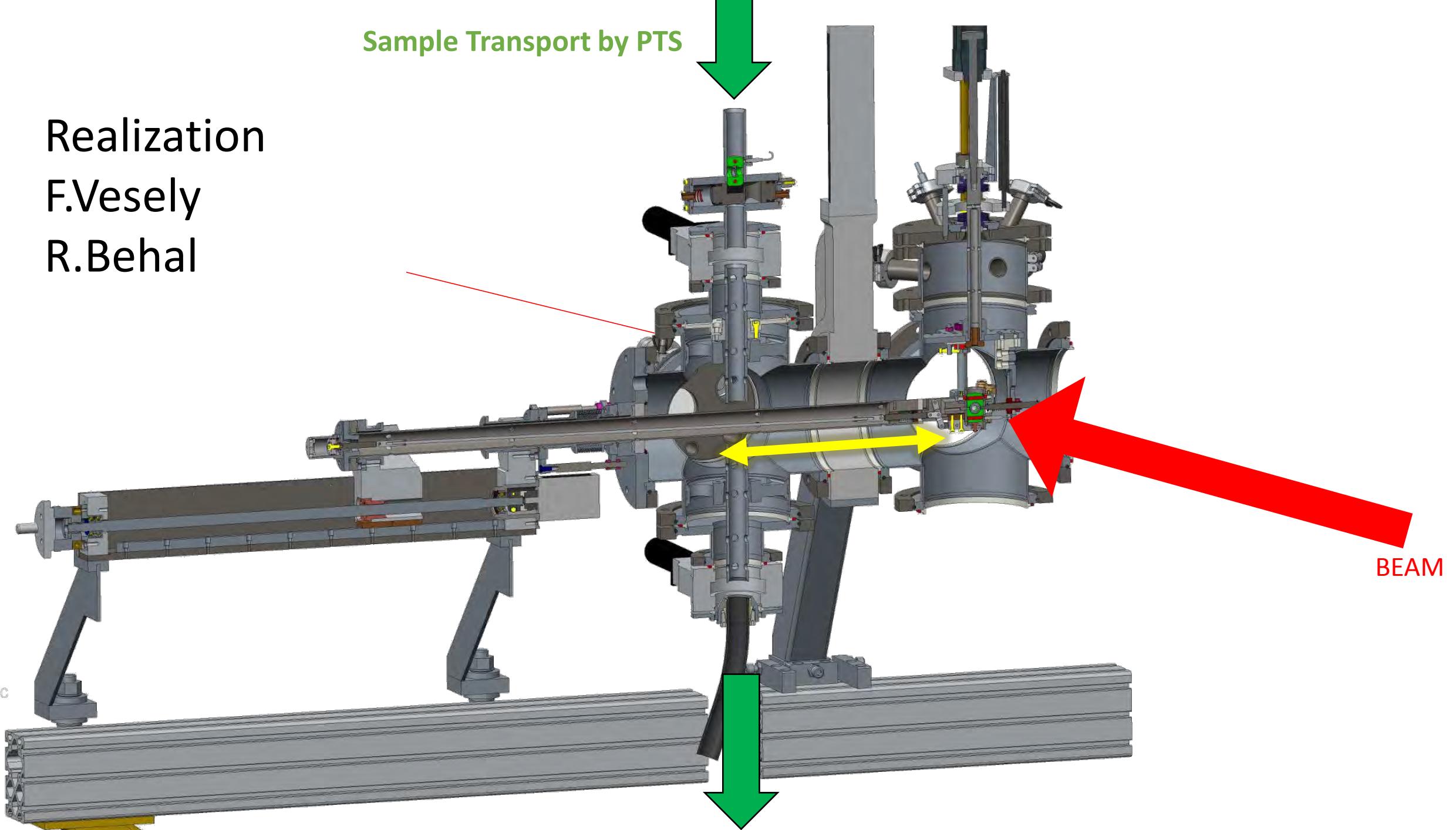
## - mechanical concept 1



See E.Simeckova presentation

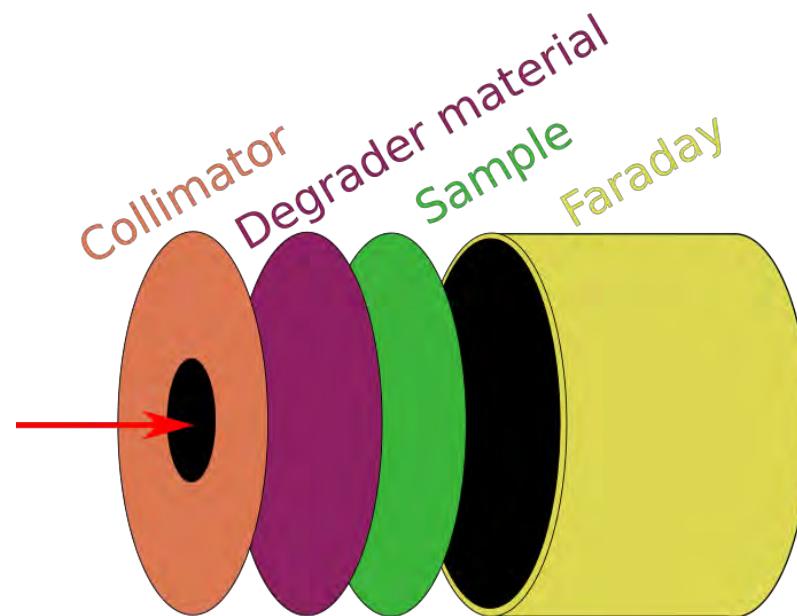
Realization  
F.Vesely  
R.Behal

Sample Transport by PTS



# Activation by charged particles

## - mechanical concept 2



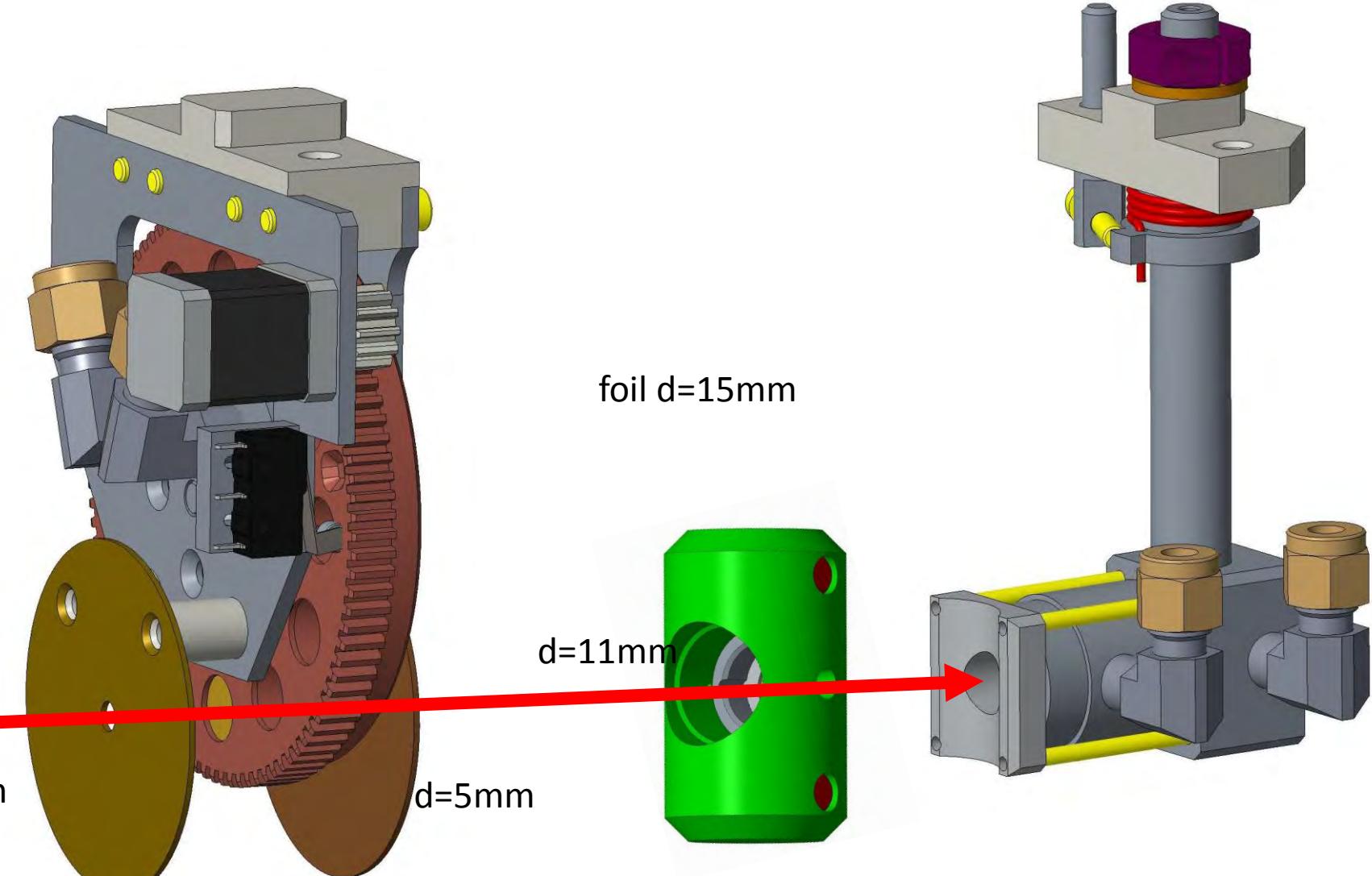
# Detail of the irradiation system – R.Behal

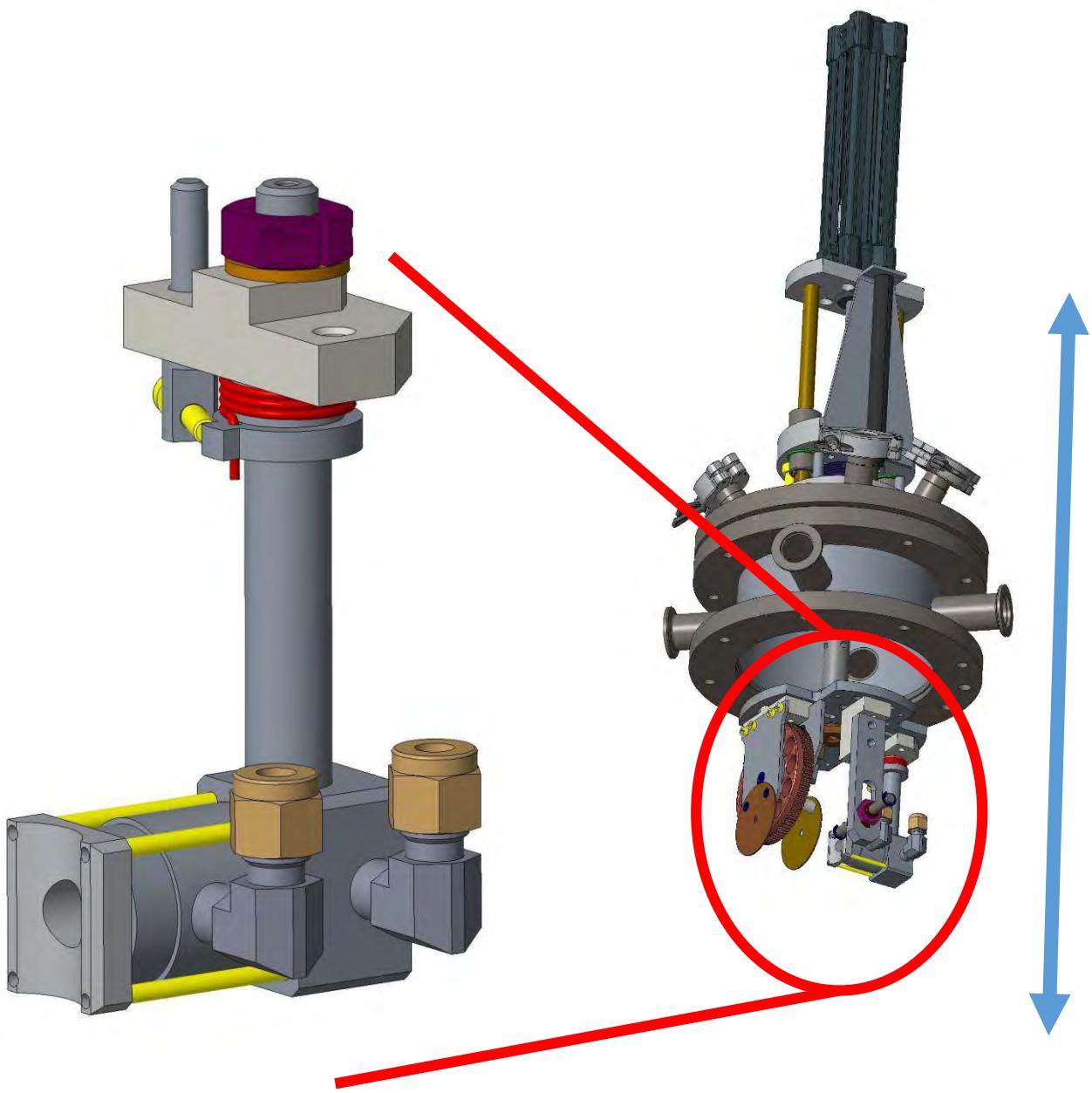
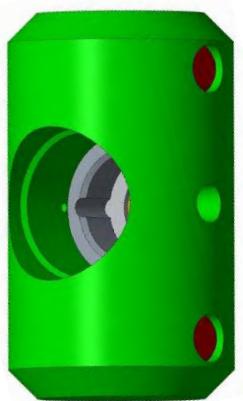
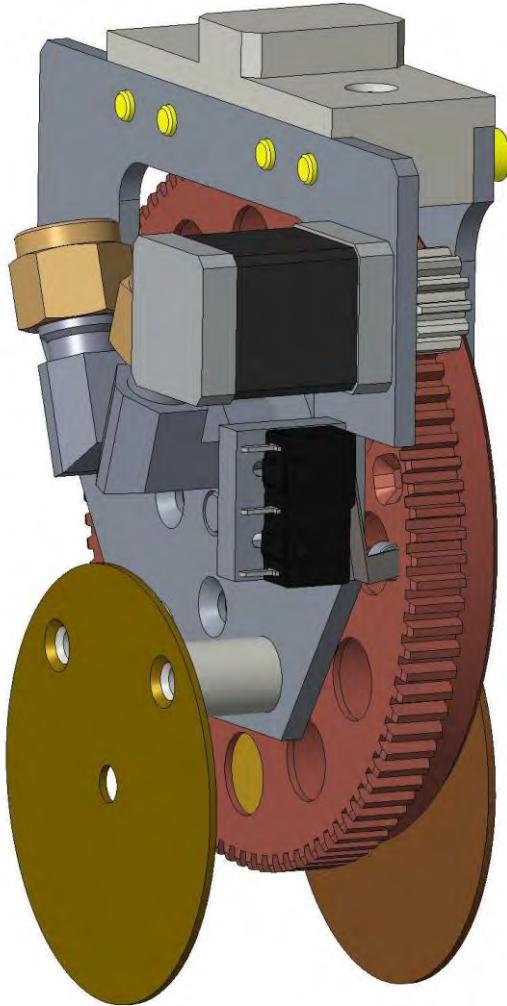
rotating **degrader** – 12 positions

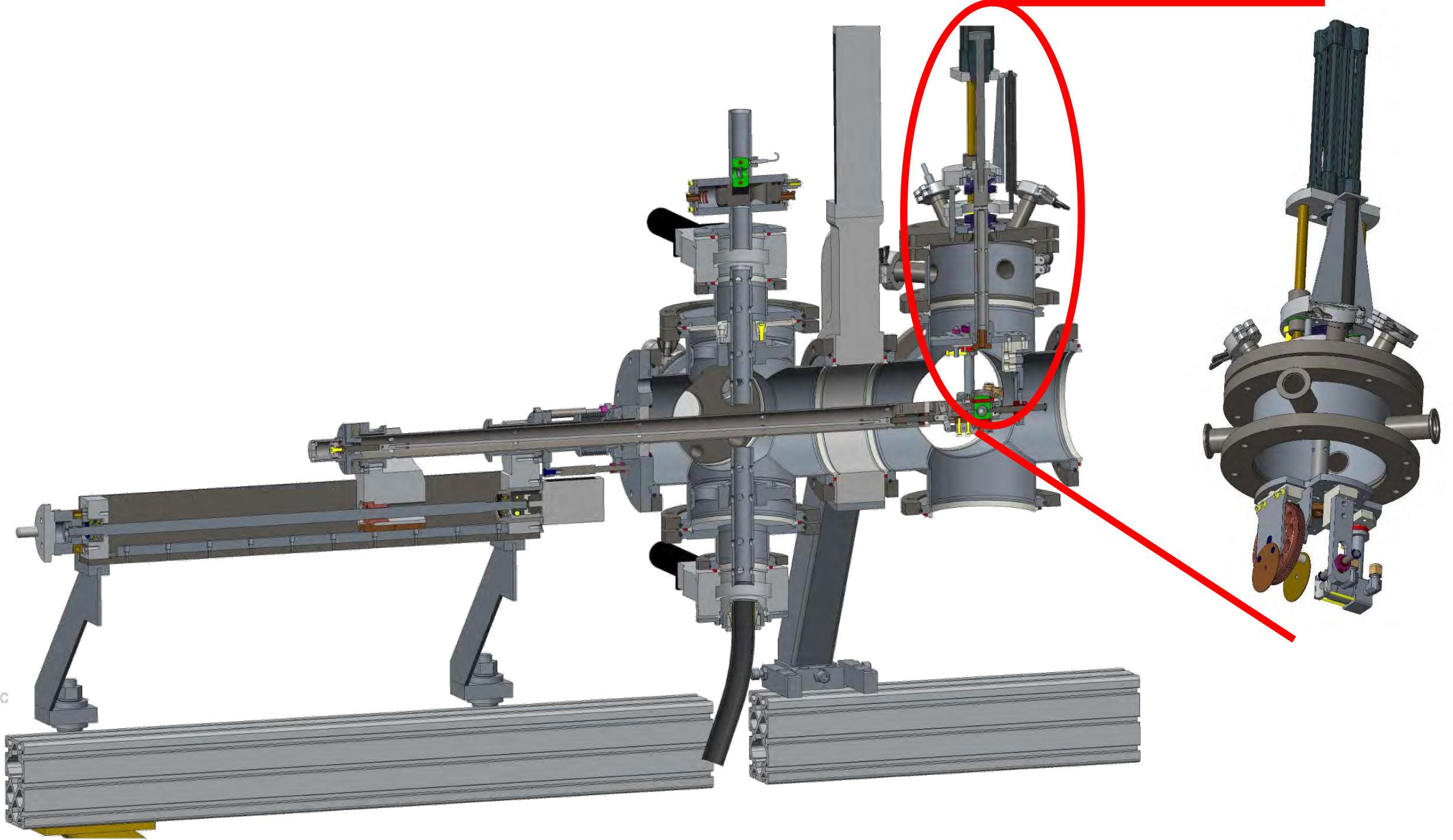
one more **electrode** between  
degrader and Faraday system

Faraday in **contact** with the **rabbit**  
- thermal  
- electrical

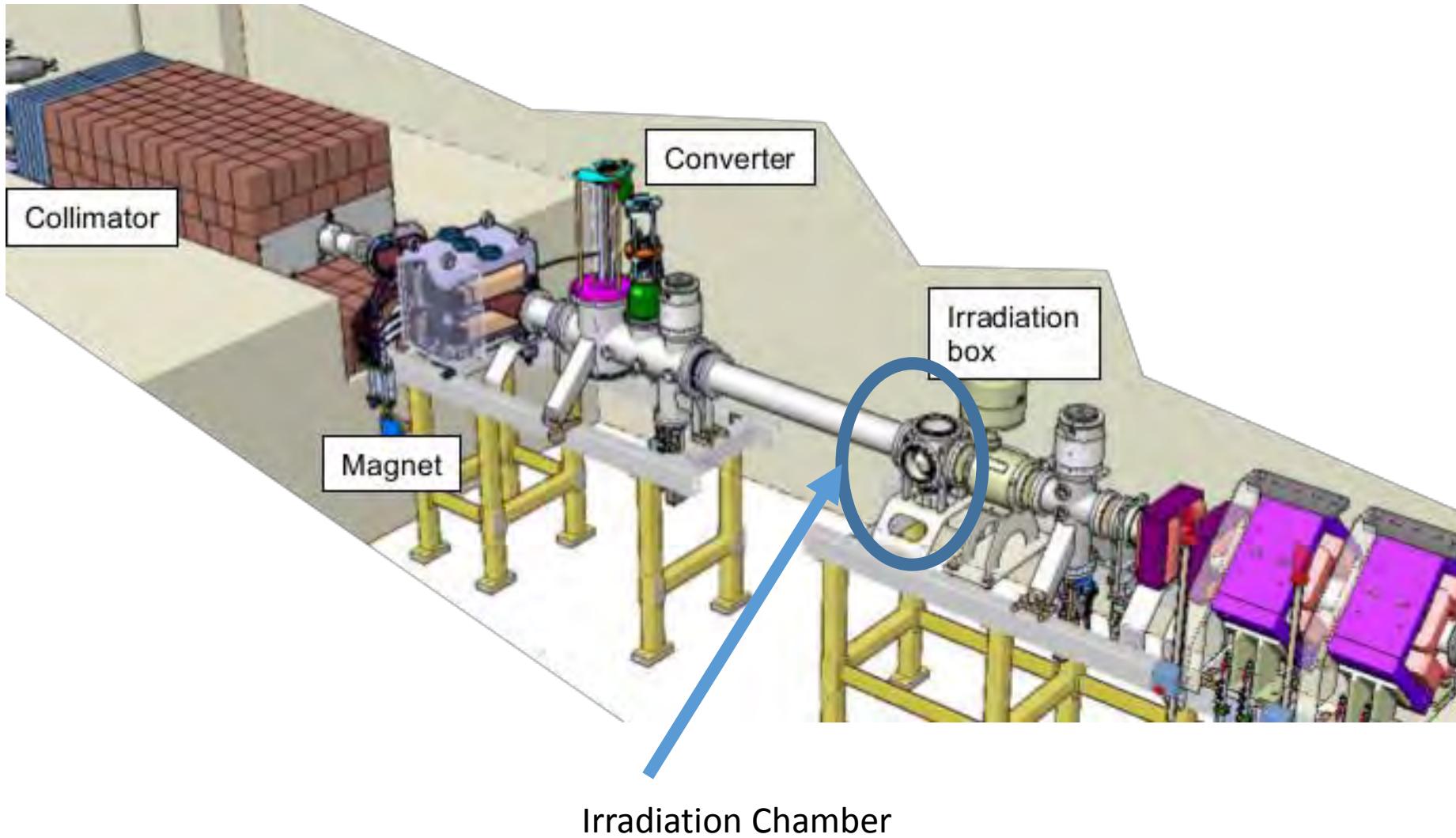
degrader and Faraday are **cooled**

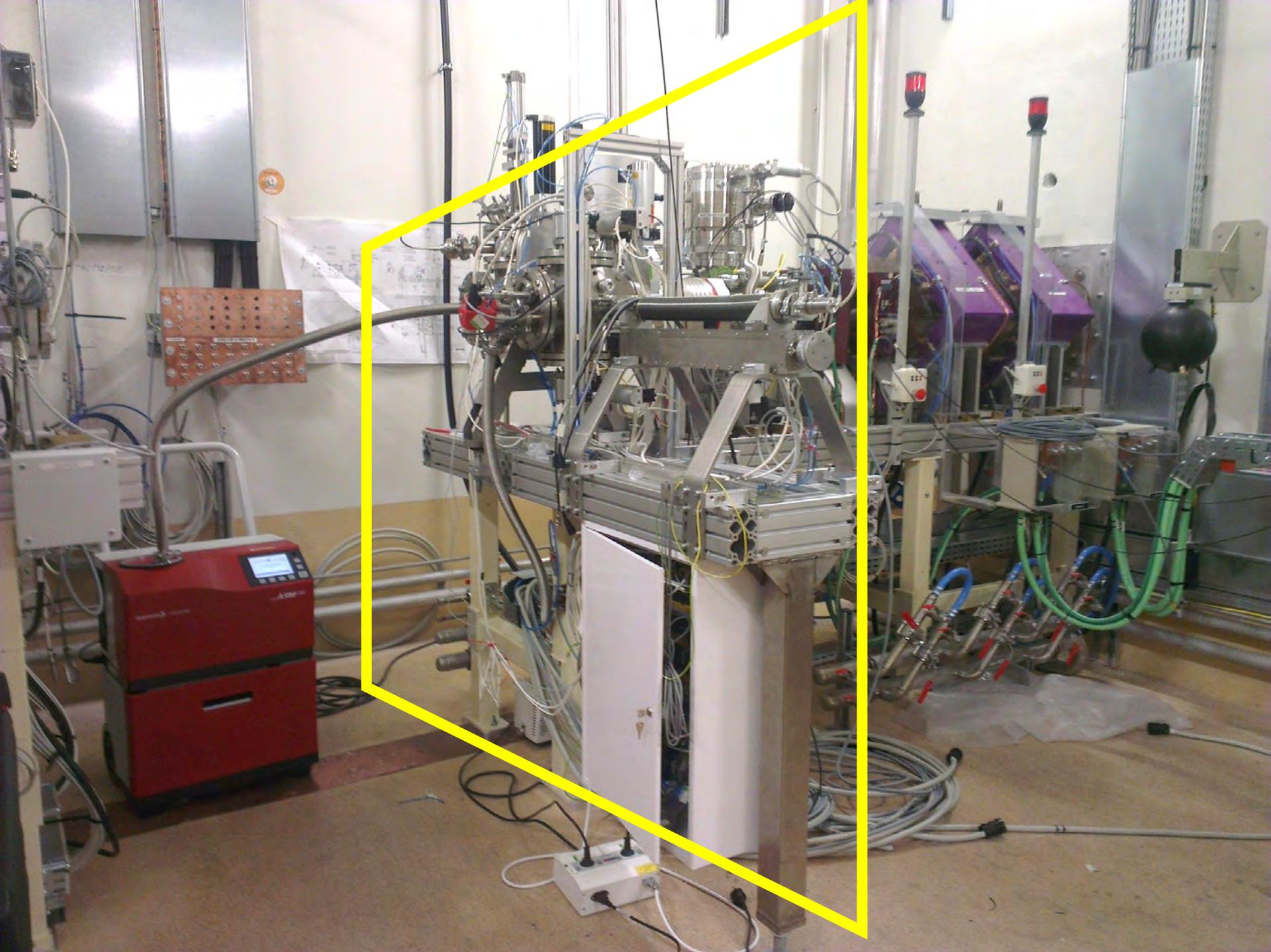




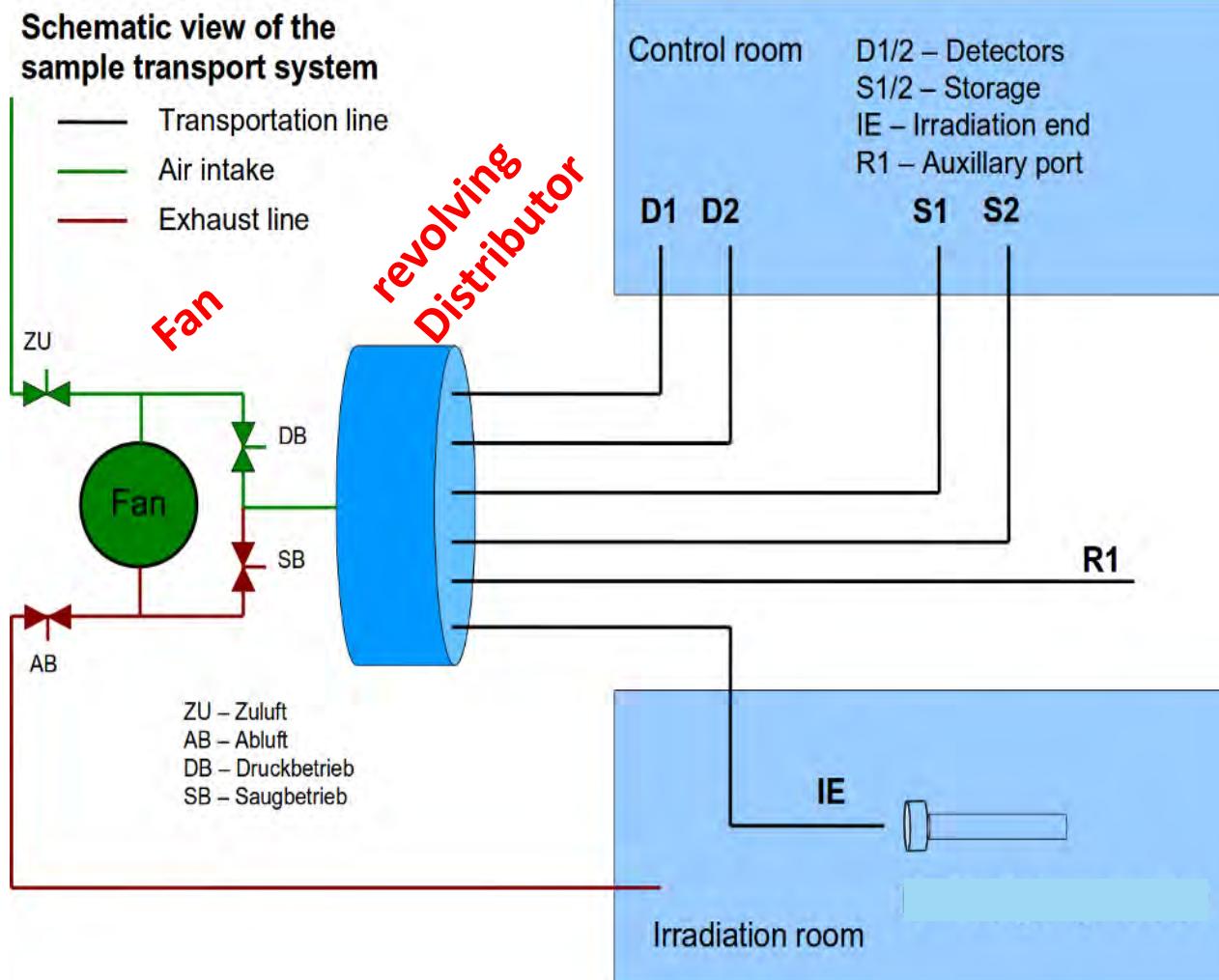


# Irradiation Chamber in SPIRAL2/NFS





# PTS – Pneumatic Transfer System from KIT for Irradiation Chamber



Delivered by A.Klix, U.Fischer  
from KIT Karlsruhe,  
based on system of TU Dresden

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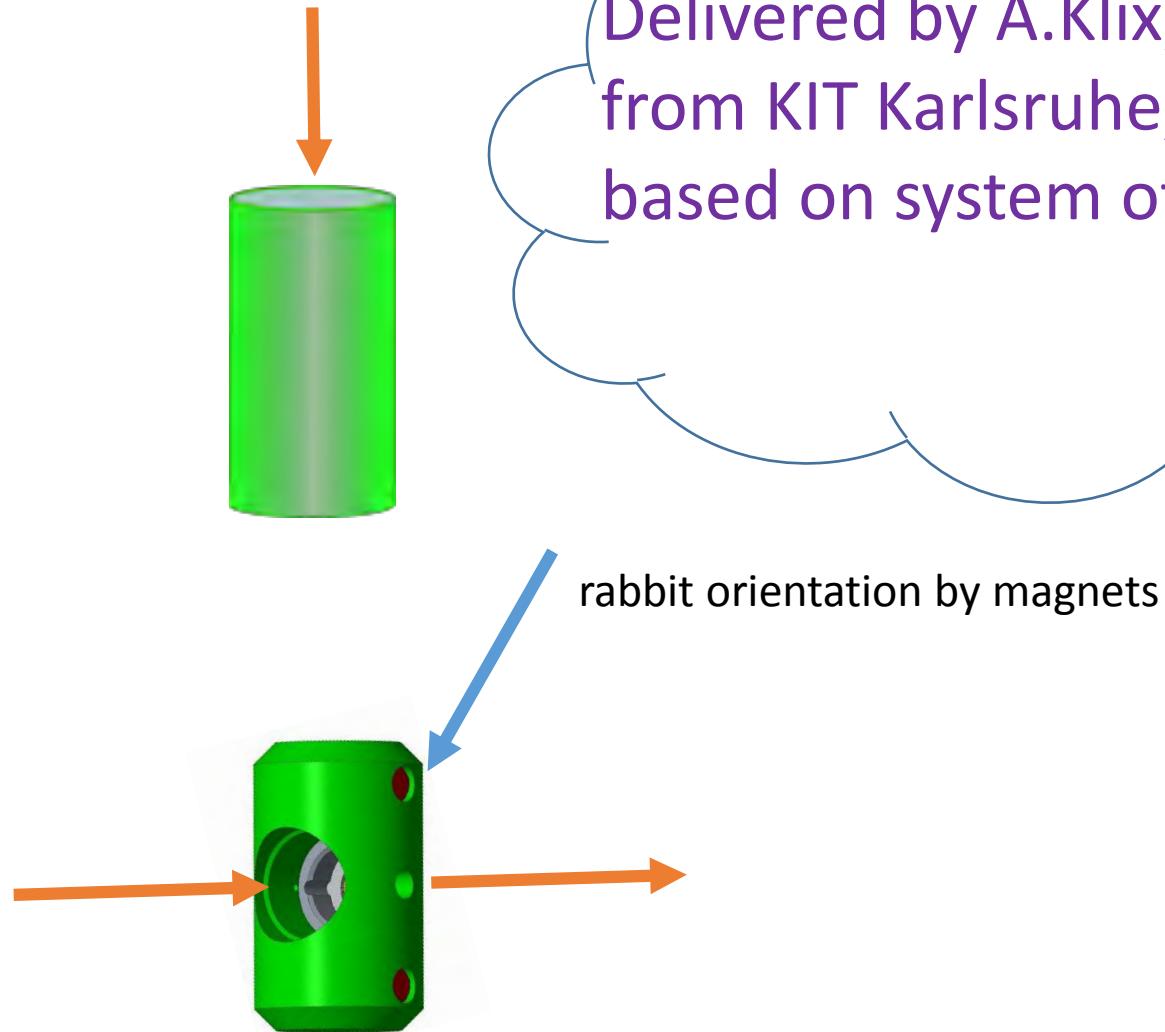
Delivered by A.Klix, U.Fischer  
from KIT Karlsruhe,  
**TU Dresden:**  
n – activation photo

# PTS – Pneumatic Transfer System from KIT for Irradiation Chamber

**Fortunate situation for NPI**

**but modifications needed:**

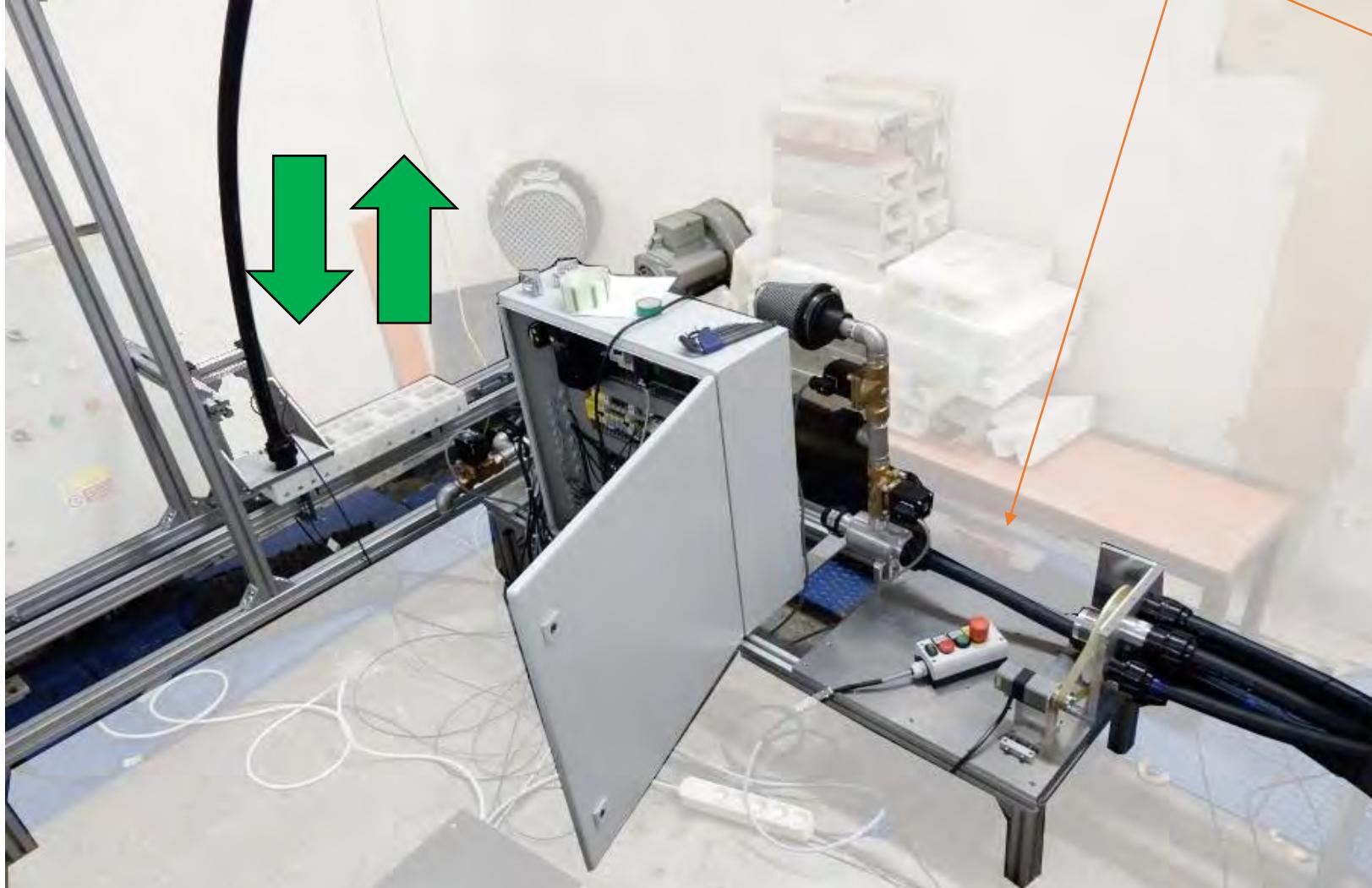
- sample – foil placement from side
- rabbit orientation at
  - HPGe station
  - Irradiation Chamber
- rabbit (Aluminium) is heavier
  - new braking system at all points
- PTS coupling to a complex control system



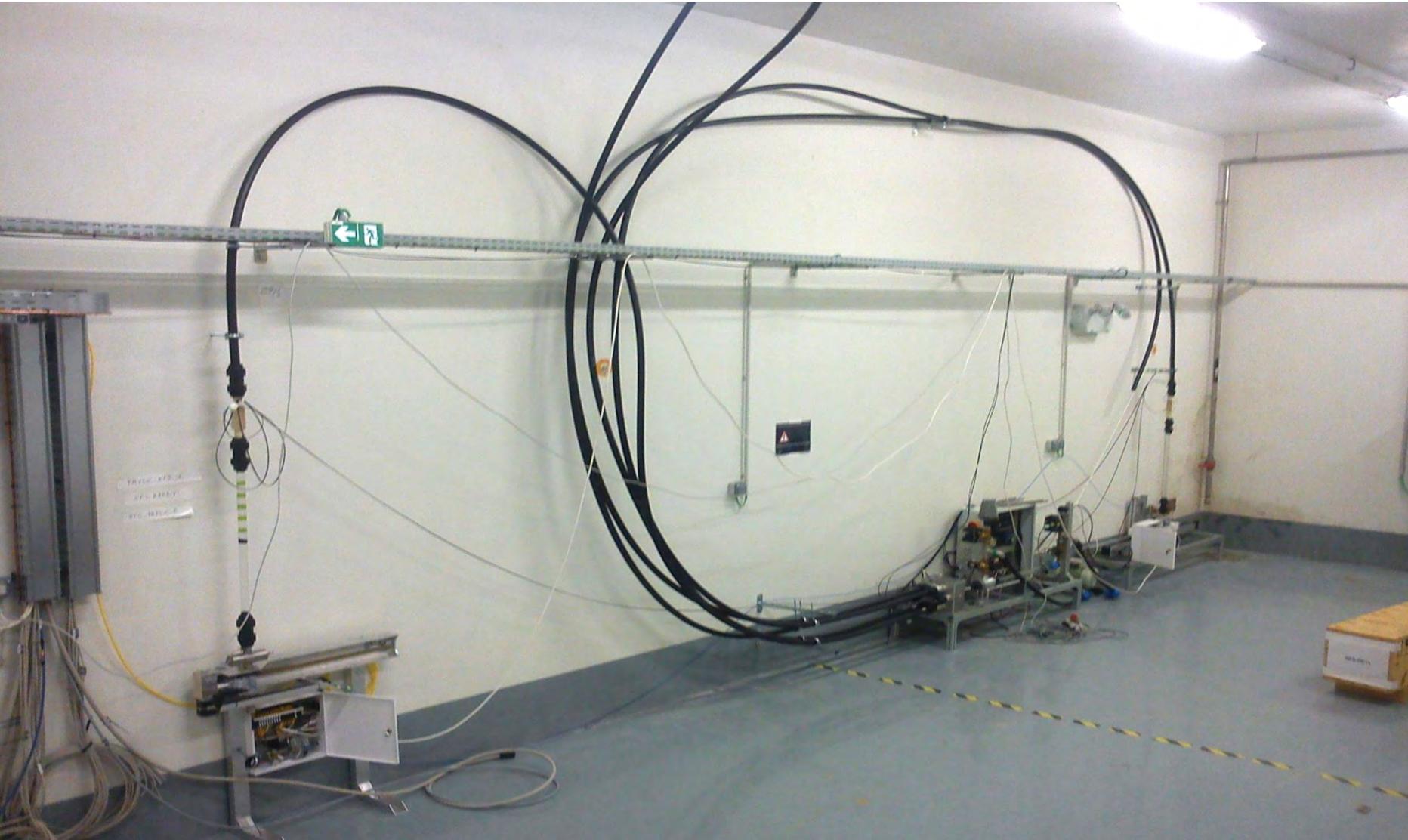
Delivered by A.Klix, U.Fischer  
from KIT Karlsruhe,  
based on system of TU Dresden

# PTS – Pneumatic Transport System from KIT

air brake system for heavy weight rabbits



# PTS in SPIRAL2/NFS - TOF hall end



## Activation by charged particles in NFS

- experimental equipment: irradiation chamber`+ PTS

## Experiments at low energies in NFS

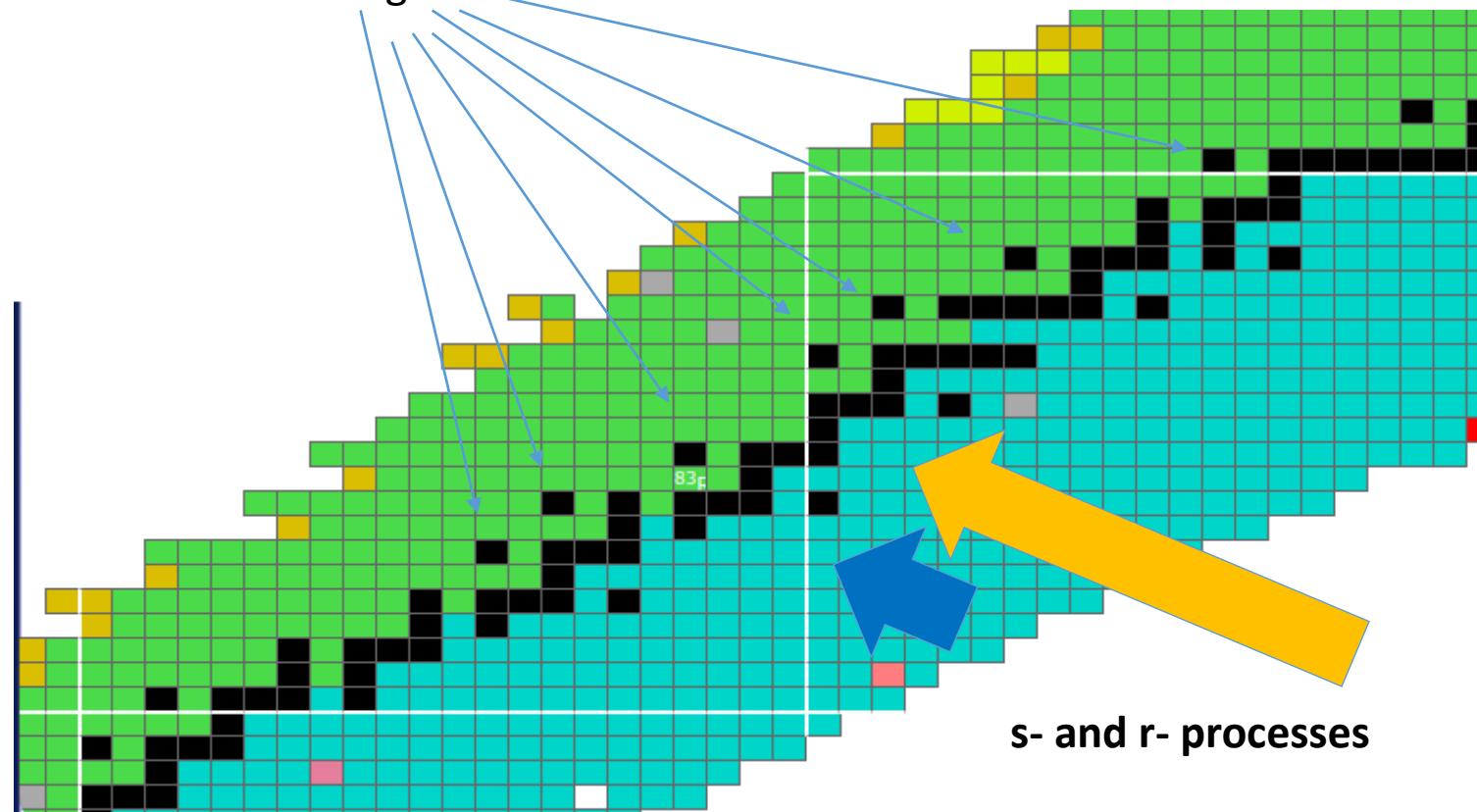
- nuclear astrophysics context

# p-process

35 nuclei

from 74Se to 196Hg

discrepancies:  
Mo, Ru, Sn, La, Gd  
underproduced



Scenario : ( $\gamma, n$ ) reactions, after ( $\gamma, p$ ) and ( $\gamma, \alpha$ ) continue

# Lol p-process : direct kinematics measurements @ SPIRAL2-NFS

Lol : B. Bastin, G. Randisi, C. Ducoin, I. Companis, S. Harissopoulos *et al.*

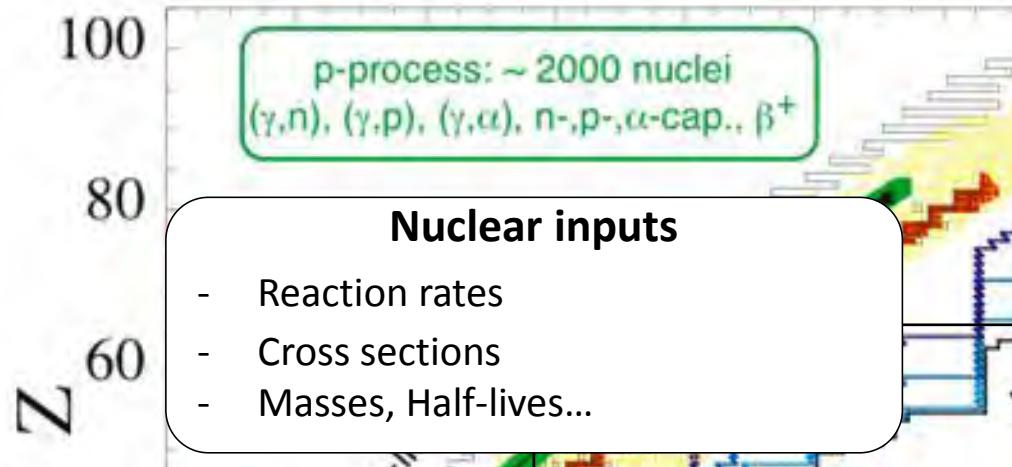
GANIL, INPL, NCSR "Demokritos", Subatech, LPC Caen, CEA-DAM, IFIN-HH, ATOMKI, University of Jyvaskyla, NPI CAS and IPN Orsay.

## Reviews :

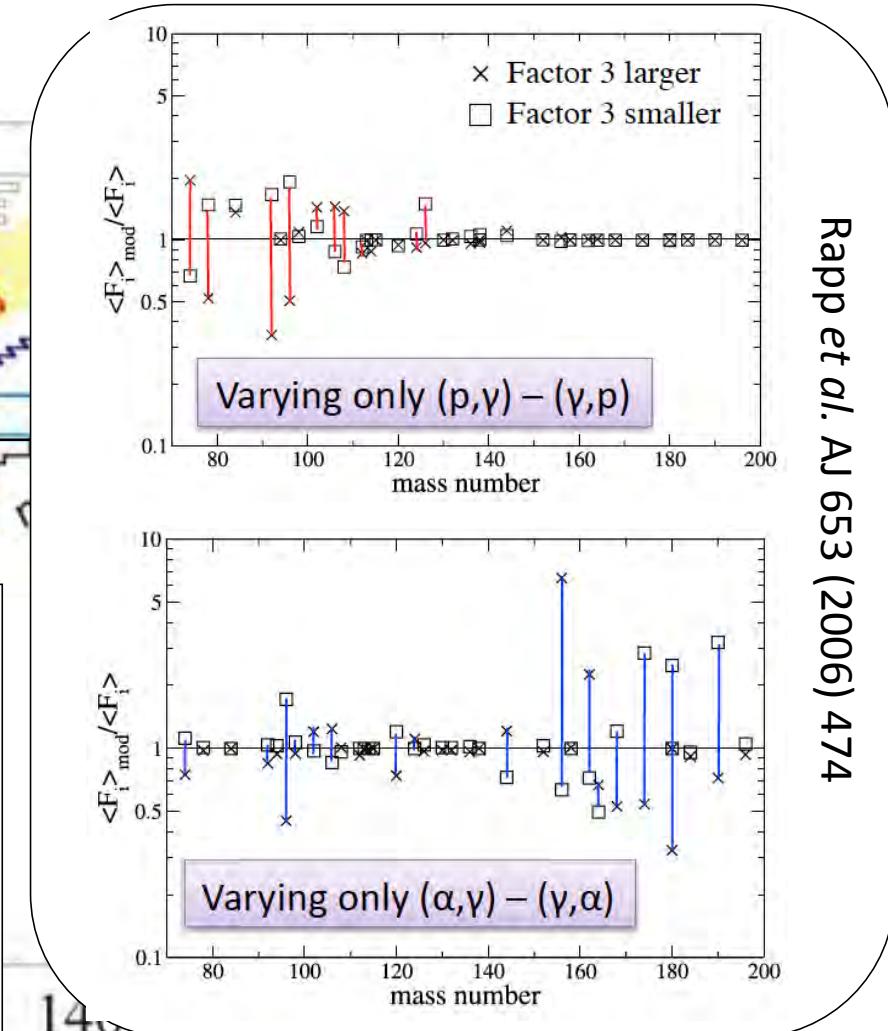
Arnould-Goriely, Physics Report 384 (2003) 1

Rauscher *et al*, Rep. Prog. Phys. 76 (2013) 066201

Rapp *et al.* AJ 653 (2006) 474



- Possible scenarios (p-process)**
- 1) O/Ne Layers of massive stars during SNII ( $1.7 \leq T_9 \leq 3.3$ )**  
Woosley & Howard 1978, Rayet 1990, 1995, Prantzos 1990, Hayakawa 2004, 2008
  - 2) Type Ia SN ( $1.5 \leq T_9 \leq 3.7$ )**  
Howard & Meyer 1992, Goriely 2001, Travaglio 2011
  - 3) Recently: vp-process, neutrino driven wind ( $1.0 \leq T_9 \leq 3.0$ )**



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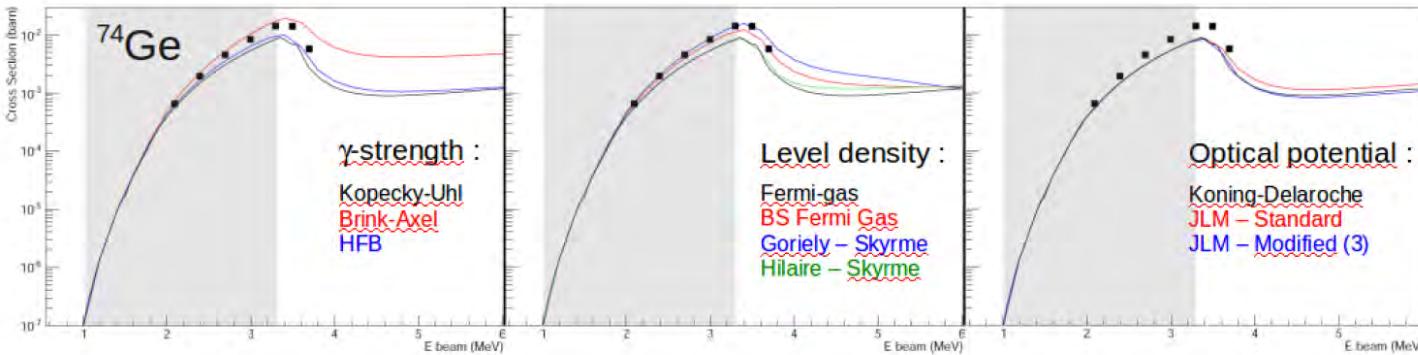


FIG. 3:  $^{74}\text{Ge}(p, \gamma)^{75}\text{As}$  cross section. Symbols : experimental data from [22]. Curves : Talys calculations, varying three input parameters : gamma strength (left), level density (middle) and optical potential (right). The shadowed region indicates the Gamow window for  $T_9 = 1.5 - 3.5$ .

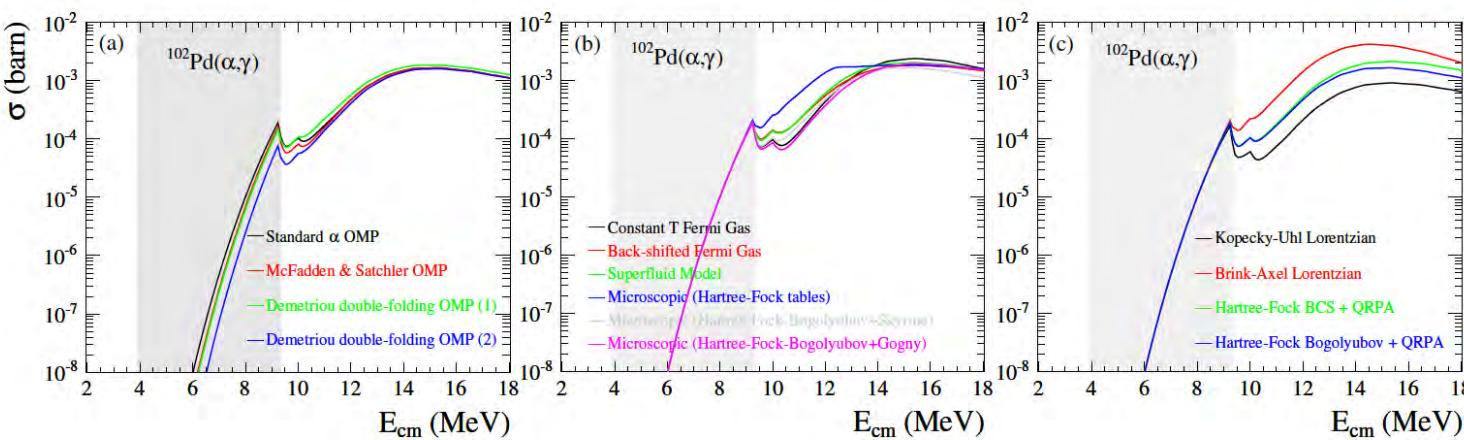


FIG. 4: Sensitivity of the  $^{102}\text{Pd}(\alpha, \gamma)$  cross section to different (a) optical model potentials [20, 21, 23], (b) nuclear level densities [24] and (c) gamma strengths [25–28]. Several phenomenological and semi-microscopic models implemented in the TALYS code [29] are compared. The shadowed area indicates the relevant energy range for ( $\alpha, \gamma$ ) reactions at  $1.5 \leq T_9 \leq 3.5$  [30].

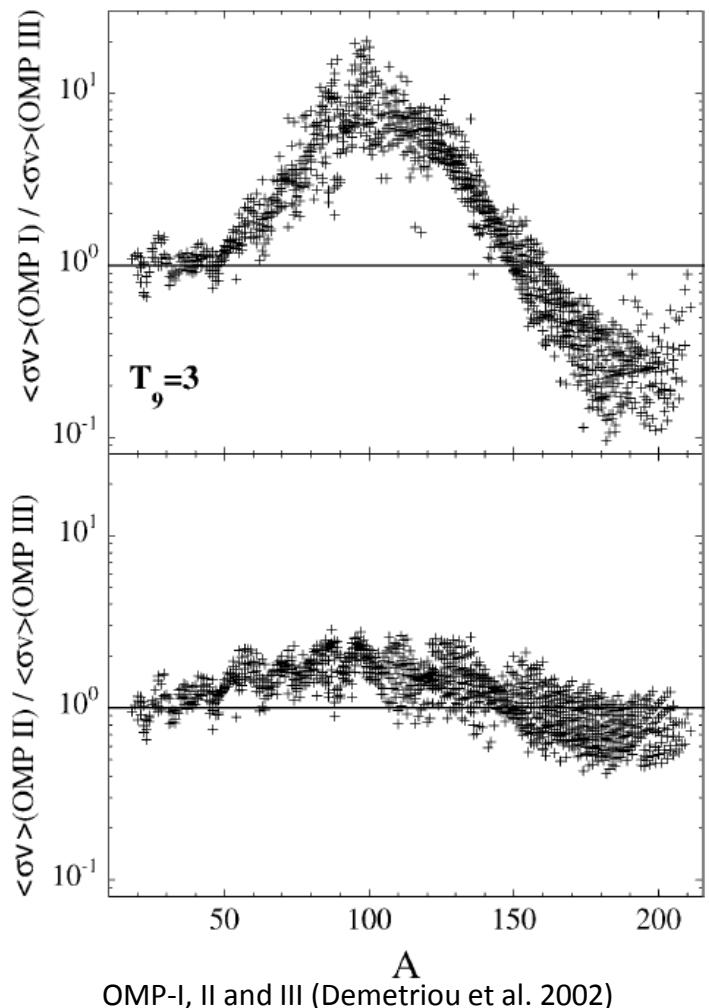
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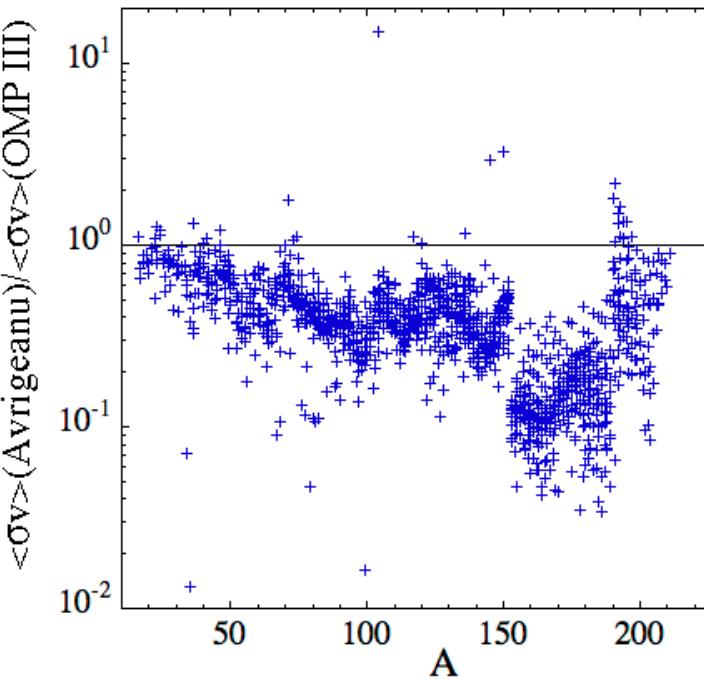
GANIL, INPL, NCSR "Demokritos", Subatech, LPC Caen, CEA-DAM, IFIN-HH, ATOMKI, University of Jyvaskyla, NPI CAS and IPN Orsay.

Different GLOBAL  $\alpha$ -OMP available: Demetriou et al. (2002); Avrigeanu et al. (2014), etc...fitted on many low-energy cross sections ( $\alpha, \gamma$ ), ( $\alpha, n$ ), ( $n, \alpha$ ), ( $\alpha, p$ ), scattering, ...

Relatively different predictions of ( $\alpha, \gamma$ ) reaction rates ( $2 \leq T_9 \leq 3$ )



$T_9 = 3$ ; stable & n-deficient nuclei with  $8 \leq Z \leq 84$



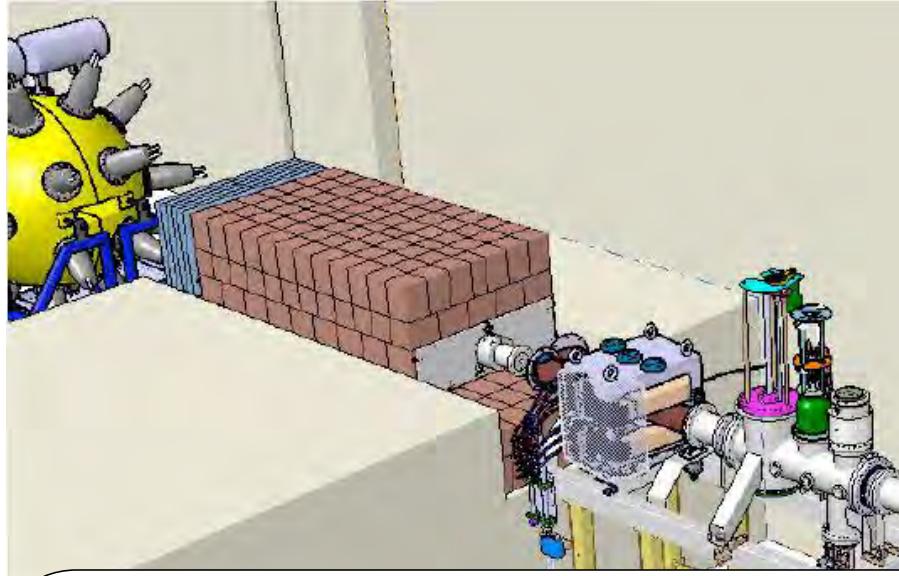
- New  $\alpha$ -OMP (S. Gorieli, V. Demetriou, C. Ducoin...)
- Experiments with  ${}^4\text{He}$  beams @ Rez

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3 experimental campagnes foreseen :  
Activation and 2 in-beam



Critical p-process Reaction Rates  
(list of day one experiments-easy cases)  
*(will be updated)*

$(p, \gamma)$	$(p, n)$	$(\alpha, \gamma)$
$^{72}\text{Ge}(p, \gamma)^{73}\text{As}$	$^{76}\text{Ge}(p, n)^{76}\text{As}$	$^{70}\text{Ge}(\alpha, \gamma)^{74}\text{Se}$
$^{74}\text{Ge}(p, \gamma)^{75}\text{As}$	$^{75}\text{As}(p, n)^{75}\text{Se}$	$^{92}\text{Mo}(\alpha, \gamma)^{96}\text{Ru}$
$^{77}\text{Br}(p, \gamma)^{78}\text{Kr}^*$	$^{85}\text{Rb}(p, n)^{85}\text{Sr}$	$^{102}\text{Pd}(\alpha, \gamma)^{106}\text{Cd}$
$^{83}\text{Rb}(p, \gamma)^{84}\text{Sr}^*$	$^{86}\text{Kr}(p, n)^{86}\text{Rb}$	$^{106}\text{Cd}(\alpha, \gamma)^{110}\text{Sn}$

note :  $(p, \gamma) : 1.5 - 5.0 \text{ MeV}$   $(\alpha, \gamma) : 3.5 - 11.0 \text{ MeV}$

Very intense low energy beams

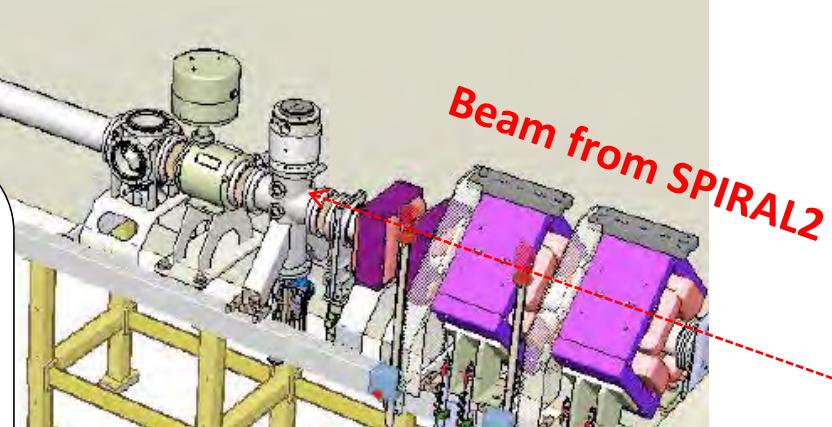
LI

A/Q=6 & SC ECR  
*(phase 1++)*

beam	P <sup>+</sup>	D <sup>+</sup>	ions	ions
Q/A	1	1/2	1/3	1/6
Max. I (mA)	5	5	1	1
Max E (MeV/A)	33	20	14.5	8
beam power (kW)	$\leq 165$	$\leq 200$	$\leq 44$	$\leq 48$

R. Ferdinand *et al.*, Proceedings of IPAC2013

Note  $E_{\min} = 0.75 \text{ MeV/u (RFQ)}$



Experiment challenge under study :  
use of radioactive targets! ★

contact person : B. Bastin

# LoI p-process : direct kinematics measurements @ SPIRAL2-NFS

LoI : B. Bastin, G. Randisi, C. Ducoin, I. Companis, S. Harissopoulos *et al.*

## measurement methods foreseen

### Activation (GANIL)

$\gamma$  decay  
Irradiation setup + OFF LINE

#### **Advantages**

Direct measurement  
Low Background  
Good resolution

#### **Constraints/Disadvantages**

Required enriched targets  
Difficult if  $T_{1/2}$  such that  $t_{\text{offline}} \geq 1 \text{ mois}$

### In-beam spectroscopy

$\gamma$  from de-excitation IN BEAM (point cible)

#### “ $\gamma$ -summing”(Demokritos)

#### **Advantages**

Covers  $4\pi$

#### **Constraints/Disadvantages**

Beam purity  
DM in case of huge count rate  
Low resolution

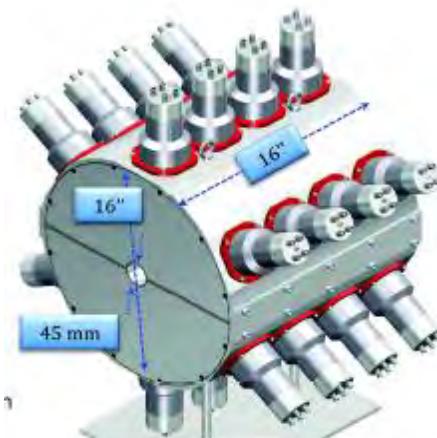
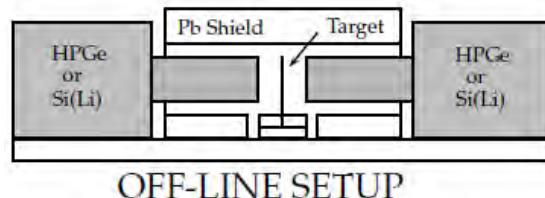
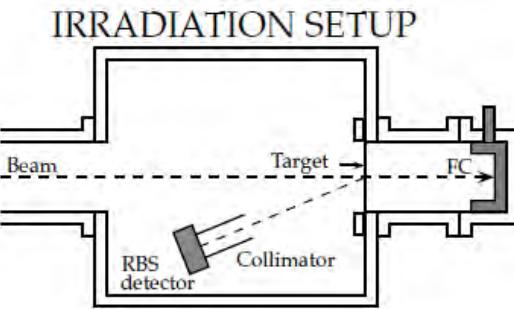
#### Angular distributions (IPNL)

#### **Advantages**

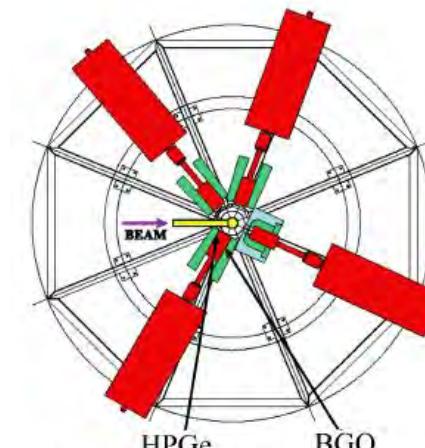
High resolution  
(compared to  $\gamma$ -summing)

#### **Constraints/Disadvantages**

Difficulty to cover  $4\pi$  (compton..)  
DM in case of huge count rate



SuN Detector (MSU)



Stuttgart HPGe Array

# Support:



EUROPEAN UNION  
European Structural and Investment Funds  
Operational Programme Research,  
Development and Education



- **LEA NuAG** - Nuclear Astrophysics and Grids
- new (2016) research infrastructure – in Czech Republic - **SPIRAL2-CZ**
- research infrastructure of NPI CAS Rez – ***open access CANAM***

**NPI CAS Rez** – Eva Simeckova, J.Mrazek, M. Stefanik, I.Sivacek, M.Majerle, M.Ansgore, O.Lebeda  
F.Vesely, R.Behal, V.Glagolev, M.Rodak, O.Jurencak, M.Gschray, J. Gabrhel...

**KIT Karlsruhe** – Axel Klix, Ulrich Fischer, Thorsten Reimann

**IFIN-HH Bucharest** – M.Avrigeanu, V.Avrigeanu, C.Costache (theory support)

**GANIL** – X.Ledoux, G. de France, J.Grynier, F. de Oliveira, B.Bastin  
J.C.Foy, Y.Georget, V.Morell, N.Menard, Gael and Alexandre Lebec ...

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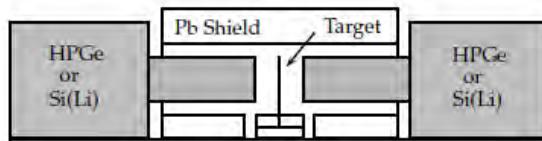
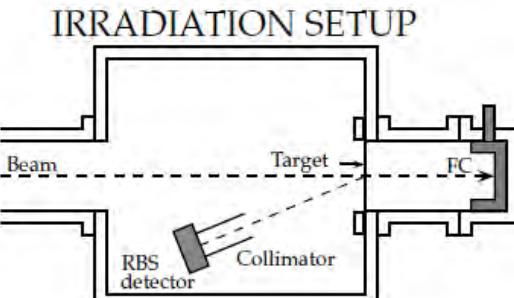
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Low Background  
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OFF-LINE SETUP

### In-beam spectroscopy

$\gamma$  from de-excitation IN BEAM (point cible)

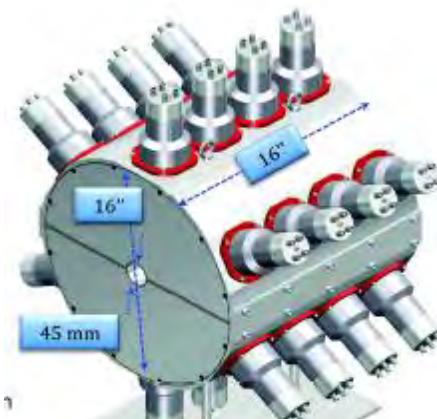
#### “ $\gamma$ -summing”(Demokritos)

#### **Advantages**

Covers  $4\pi$

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SuN Detector (MSU)

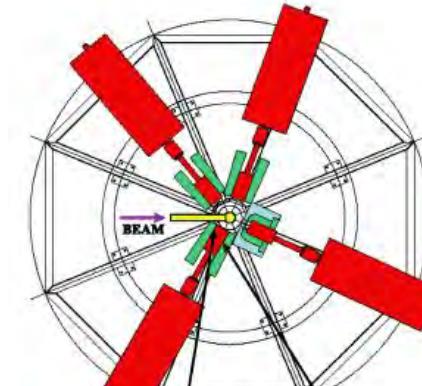
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#### **Advantages**

High resolution  
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#### **Constraints/Disadvantages**

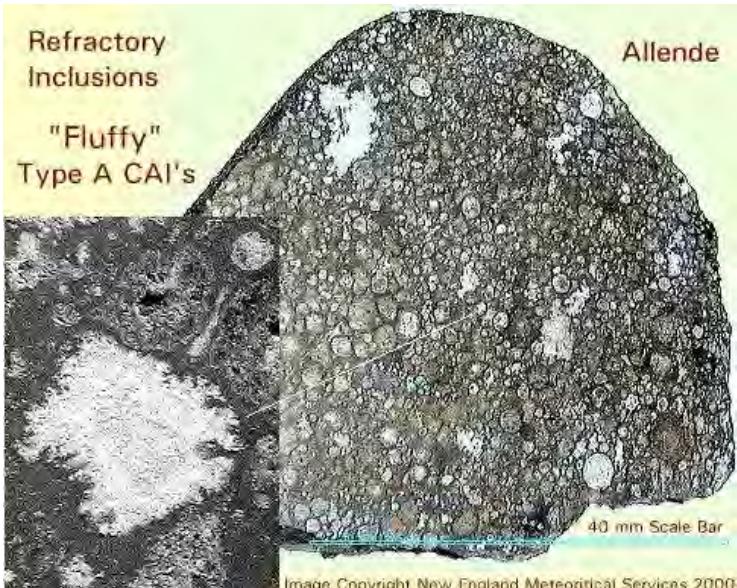
Difficulty to cover  $4\pi$  (compton..)  
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Stuttgart HPGe Array

Courtesy of Beyhan Bastin, GANIL

# Isotopic abundances



Carbonaceous chondrites

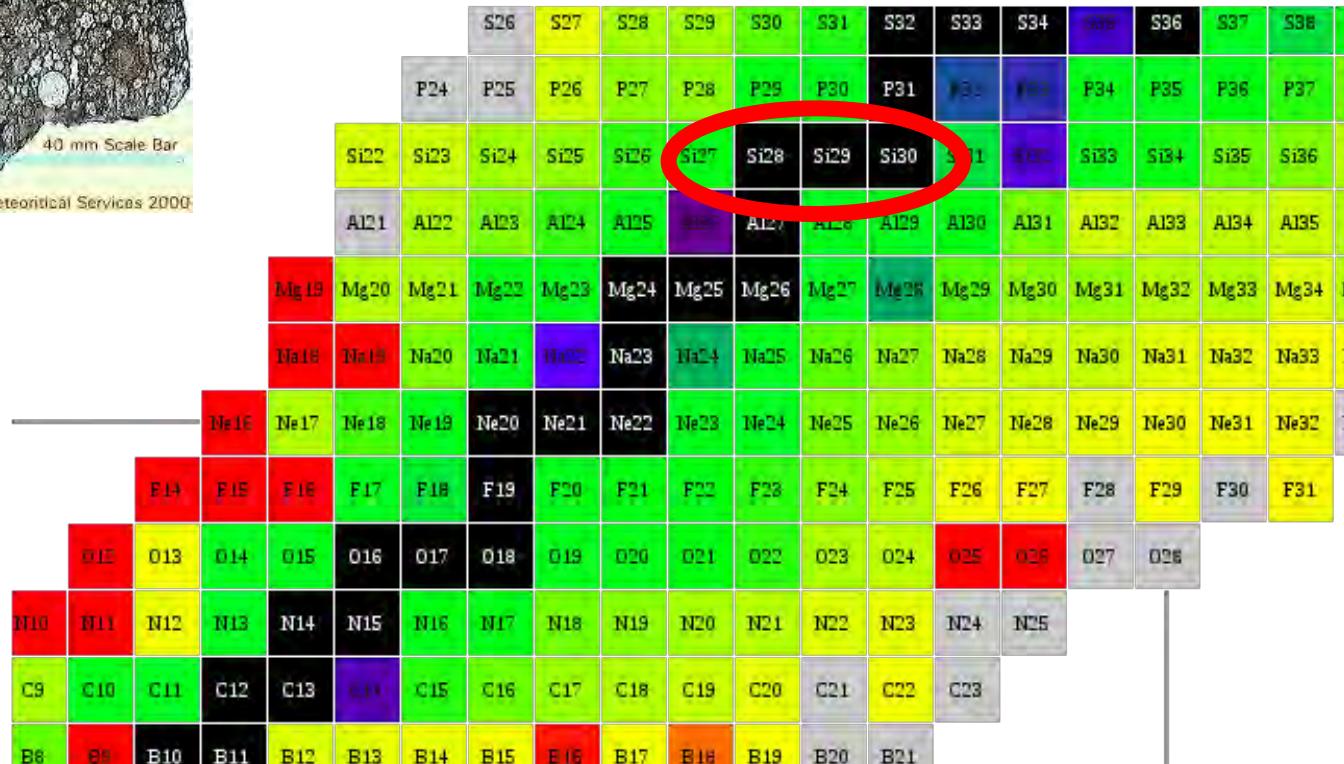


Solar system contains material from multiple dead stars  
in Solar system – isotopic ratios are equilibrated

due to mixing during formation

Inclusions in meteorites are of interest

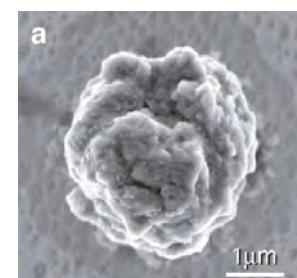
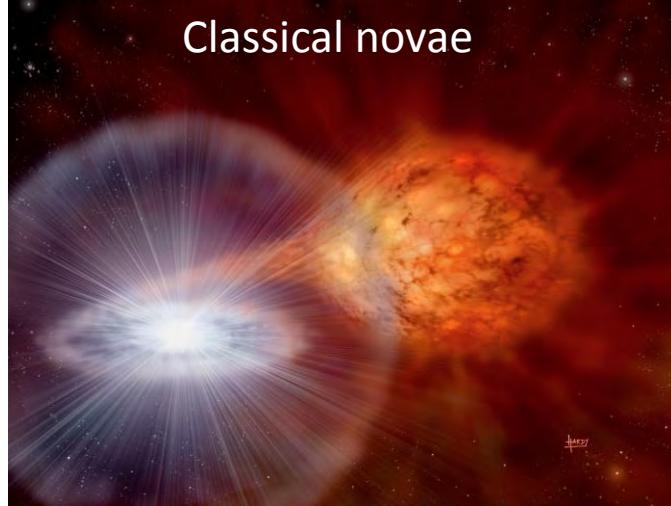
- old ( not undergone a complete mixing )
- extrasolar – information on variations



# E719 : Precise direct measurements of the key $^{28}\text{Si}(\text{p},\gamma)^{29}\text{P}$ and $^{29}\text{Si}(\text{p},\gamma)^{30}\text{P}$ reaction rates to understand the origin of presolar nova grains

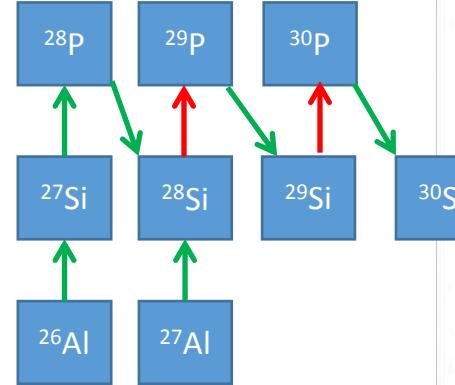
(F. Boulay, B. Bastin, J. Mrazek)

GANIL, CEA-DAM, NPI CAS, LPC Caen, IPN Orsay, CSNSM, IPN Lyon, JYFL, Instituto de Fisica Corpuscular (Valencia), NCSR "Demokritos" and Subatech



Presolar SiC grain

7 micrometeorites collected on Earth are identified coming from nova explosion

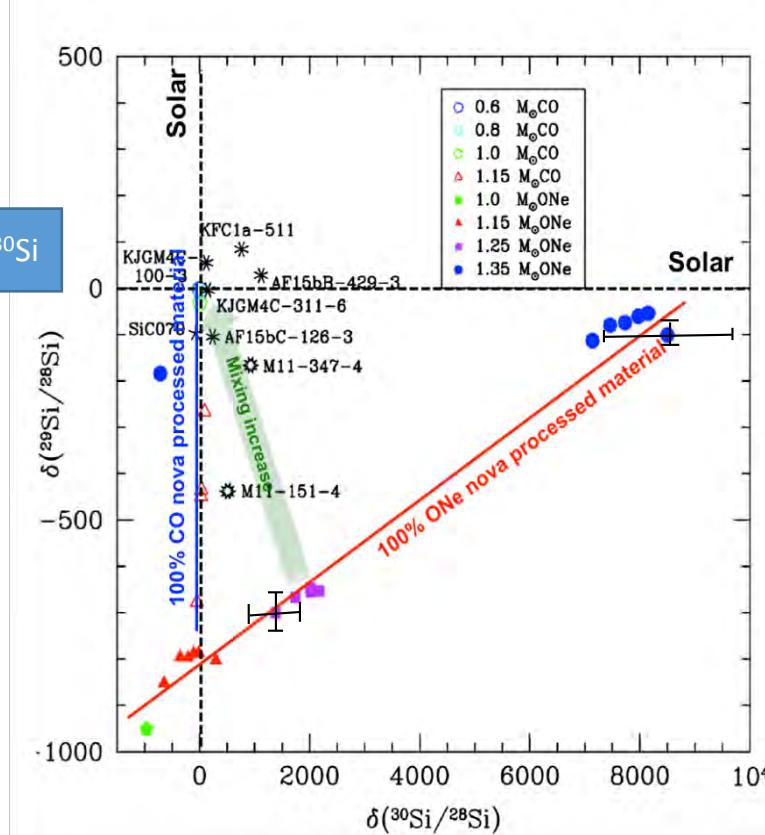


Destruction reactions  
 $^{28}\text{Si}(\text{p}, \gamma)^{29}\text{P}$ ,  $^{29}\text{Si}(\text{p}, \gamma)^{30}\text{P}$

Necessity to constrain the reaction rates  $^{28}\text{Si}(\text{p}, \gamma)^{29}\text{P}$  and  $^{29}\text{Si}(\text{p}, \gamma)^{30}\text{P}$  which have currently 21 % and 30 % uncertainties.

Aim : reduce the uncertainties on the reaction rates  $^{28}\text{Si}(\text{p}, \gamma)^{29}\text{P}$  and  $^{29}\text{Si}(\text{p}, \gamma)^{30}\text{P}$  as much as possible in the Gamow window (GW) (60 ->560 keV).

Courtesy of Beyhan Bastin, GANIL



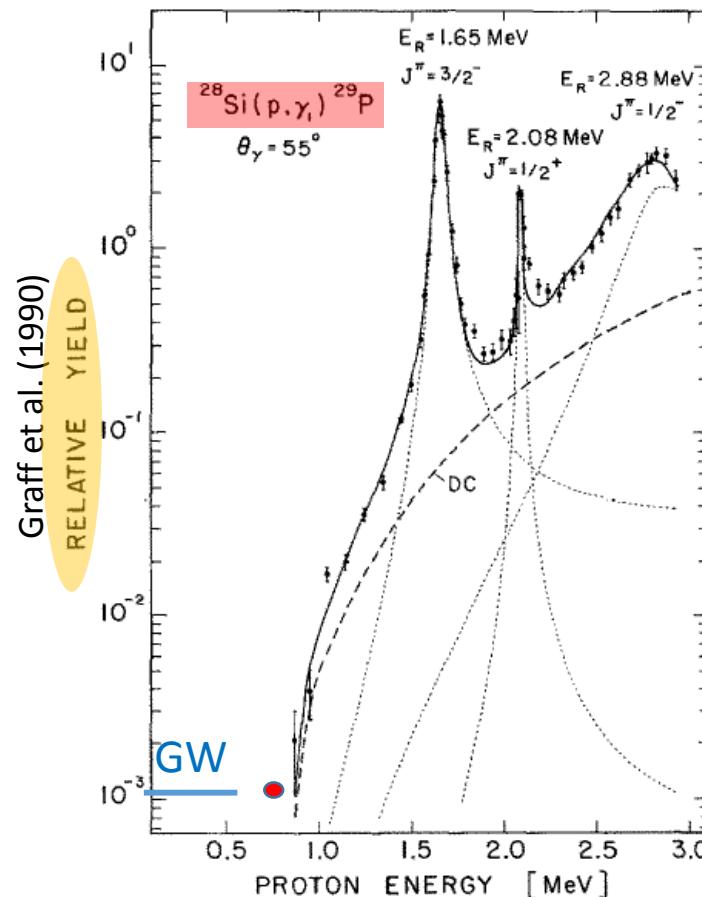
# E719 : Precise direct measurements of the key $^{28}\text{Si}(\text{p},\gamma)^{29}\text{P}$ and $^{29}\text{Si}(\text{p},\gamma)^{30}\text{P}$ reaction rates to understand the origin of presolar nova grains

(F. Boulay, B. Bastin, J. Mrazek)

GANIL, CEA-DAM, NPI CAS, LPC Caen, IPN Orsay, CSNSM, IPN Lyon, JYFL, Instituto de Fisica Corpuscular (Valencia), NCSR "Demokritos" and Subatech

## State of art

- No measurement from 1990
- Current measurement at low energy at LENA facility for  $^{29}\text{Si}(\text{p},\gamma)^{30}\text{P}$  => International competition.



$$Y = \frac{N_r}{\eta(E) N_p} = \frac{M_p + M_t}{M_t} \frac{\lambda_r^2}{2} \frac{\omega\gamma}{\epsilon_r(E)}$$

	Graff et al. (1990) @653 keV
Efficiency detection ( $\eta$ )	5 %
Statistics (Nr)	30 %
Current measurement (Np)	5 % <span style="color:red">→</span> Can be improved
Total Yield	31 %
Beam energy	0.5 %
Stopping power	10 %
Total gamma strength	31 %

Several measurements at different energies in the most intense proton beam facilities in the world has to be done

SPIRAL 2 facility will provide the most intense proton beam at 0.733 MeV in Europe

Courtesy of Beyhan Bastin, GANIL

# E719 : Precise direct measurements of the key $^{28}\text{Si}(\text{p},\gamma)^{29}\text{P}$ and $^{29}\text{Si}(\text{p},\gamma)^{30}\text{P}$ reaction rates to understand the origin of presolar nova grains

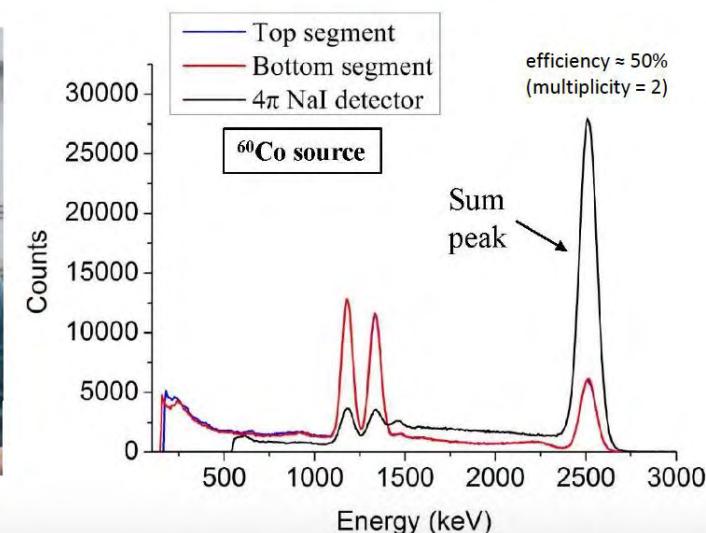
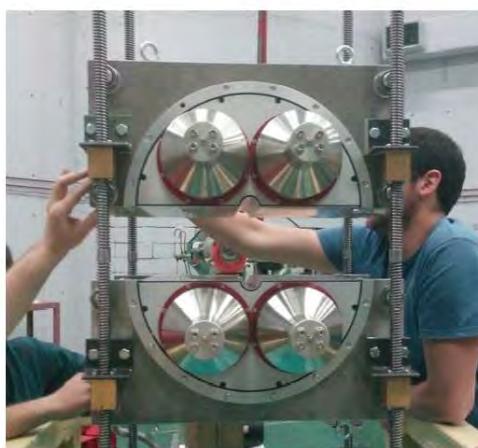
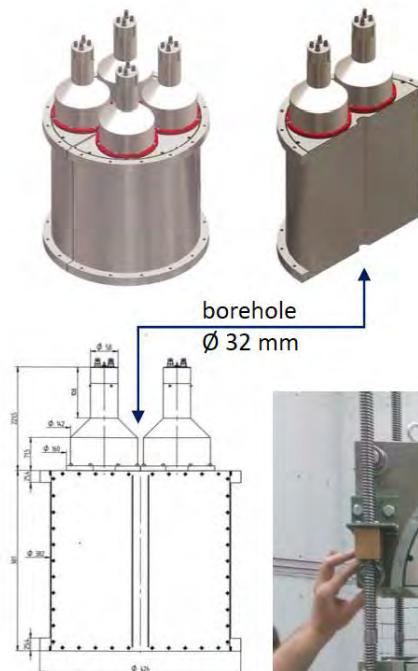
(F. Boulay, B. Bastin, J. Mrazek)

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NeoPtolemos: The  $4\pi$   $\gamma$ -summing detector at the TANDEM Accelerator Lab. of "Demokritos".

A 14''x14'' cylindrically-shaped NaI(Tl) crystal segmented in two parts



Courtesy of Beyhan Bastin, GANIL

Sotirios Harissopoulos

# Support:



EUROPEAN UNION  
European Structural and Investment Funds  
Operational Programme Research,  
Development and Education



- **LEA NuAG** - Nuclear Astrophysics and Grids
- new (2016) research infrastructure – in Czech Republic - **SPIRAL2-CZ**
- research infrastructure of NPI CAS Rez – ***open access CANAM***

**NPI CAS Rez** – Eva Simeckova, J.Mrazek, M. Stefanik, I.Sivacek, M.Majerle, M.Ansgore, O.Lebeda  
F.Vesely, R.Behal, V.Glagolev, M.Rodak, O.Jurencak, M.Gschray, J. Gabrhel...

**KIT Karlsruhe** – Axel Klix, Ulrich Fischer, Thorsten Reimann

**IFIN-HH Bucharest** – M.Avrigeanu, V.Avrigeanu, C.Costache (theory support)

**GANIL** – X.Ledoux, G. de France, J.Grynier, F. de Oliveira, B.Bastin  
J.C.Foy, Y.Georget, V.Morell, N.Menard, Gael and Alexandre Lebec ...

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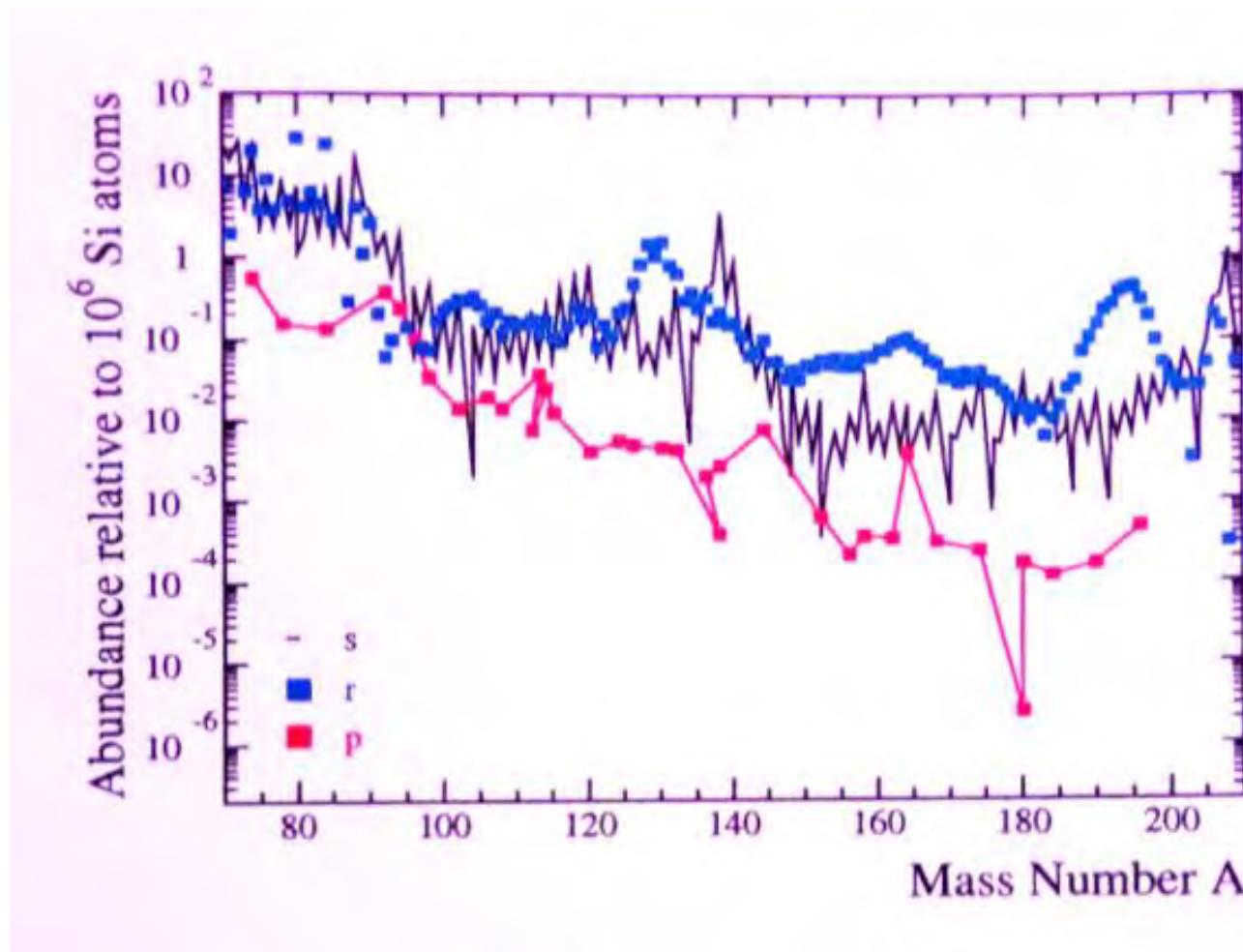
end. thank you

# p-process

35 nuclei

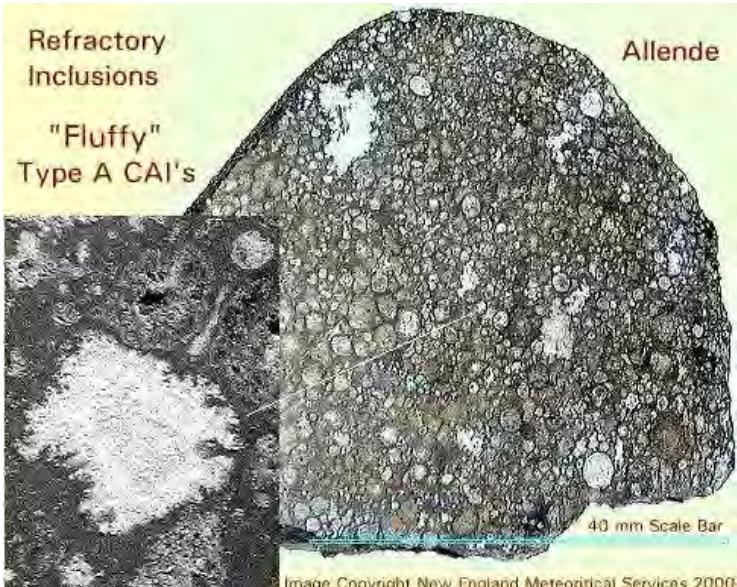
from 74Se to 196Hg

discrepancies:  
Mo, Ru, Sn, La, Gd  
underproduced

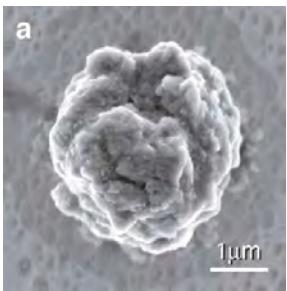


Scenario : ( $\gamma, n$ ) reactions, after ( $\gamma, p$ ) and ( $\gamma, \alpha$ ) continue

# Isotopic abundances



Carbonaceous chondrites



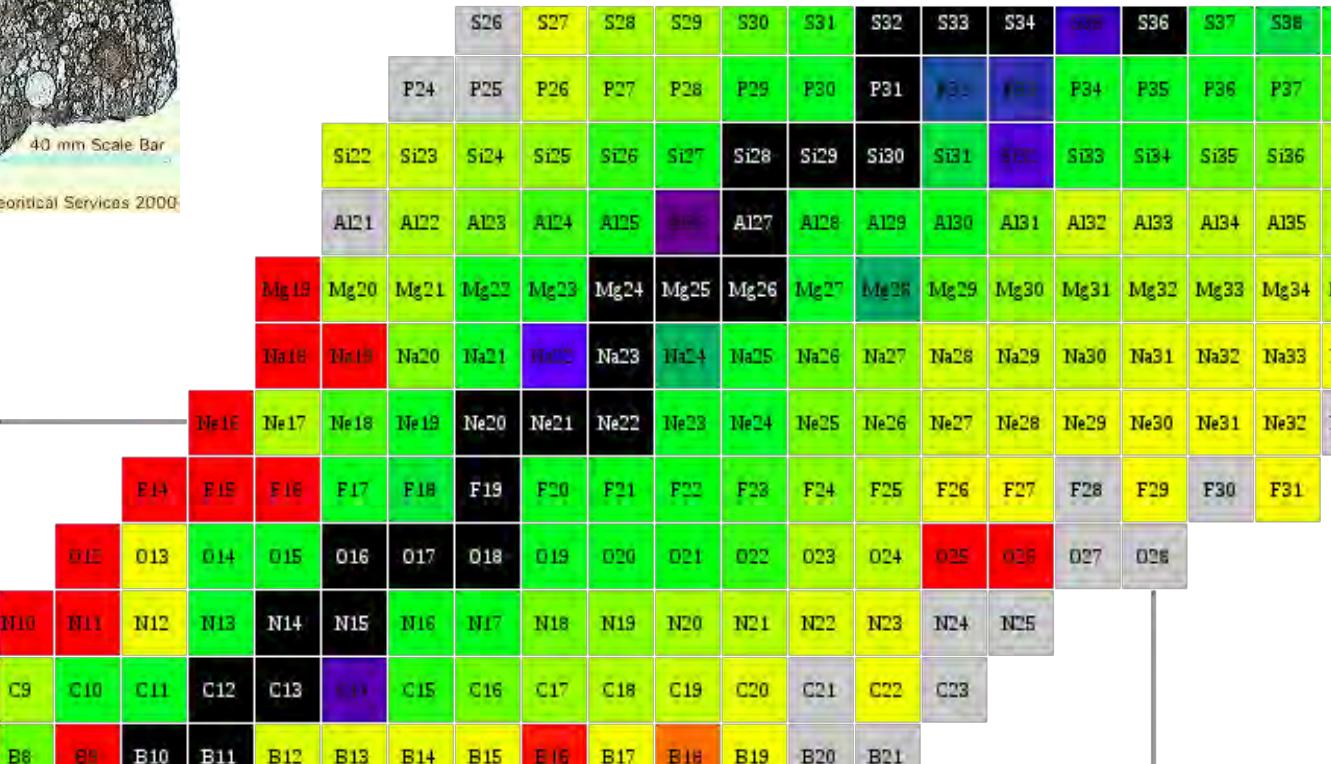
SiC

**Solar system contains material from multiple dead stars**  
in Solar system – isotopic ratios are equilibrated

due to mixing during formation

Inclusions in meteorites are of interest

- old ( not undergone a complete mixing )
- extrasolar – information on variations



# Aktivace nabitými částicemi na NFS

responsible Eva Šimečková

První experiment byl schválen na Program Advisory Committee

## ***Excitační funkce krátko-žijících izotopů na $^{nat}$ Fe***

- krátkožijící izotopy a izomery Fe, Co, Mn, Cr
- možná vůbec první experiment na SPIRAL2

**SPIRAL2-CZ-OP**

další plány: alfa + Zn, alfa +  $^{124}$ I

# Radioizotopy – výzkum pro medicínu

responsible Ondřej Lebeda

Nová tematika v GANIL/SPIRAL2 –

Odpověď na společenský požadave  
Normandie)

Diagnostika a therapeutika = therapy

Značené molekuly  $^{209}\text{Bi}(\alpha, 2n) ^{211}\text{At}$ , kandidát na theranostikum  $^{68}\text{Ga}$

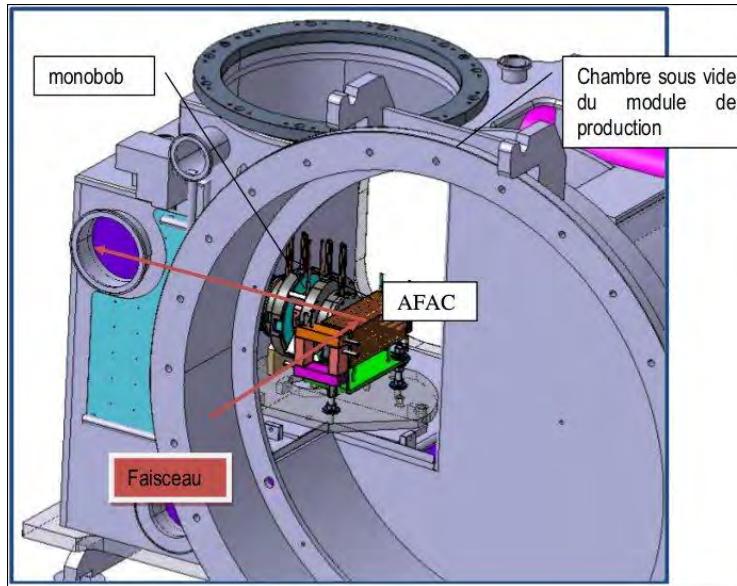
ORF ÚJF AVČR – dlouholeté zkušenosti s výzkumem

Investice + výzkum SPIRAL2-CZ-OP

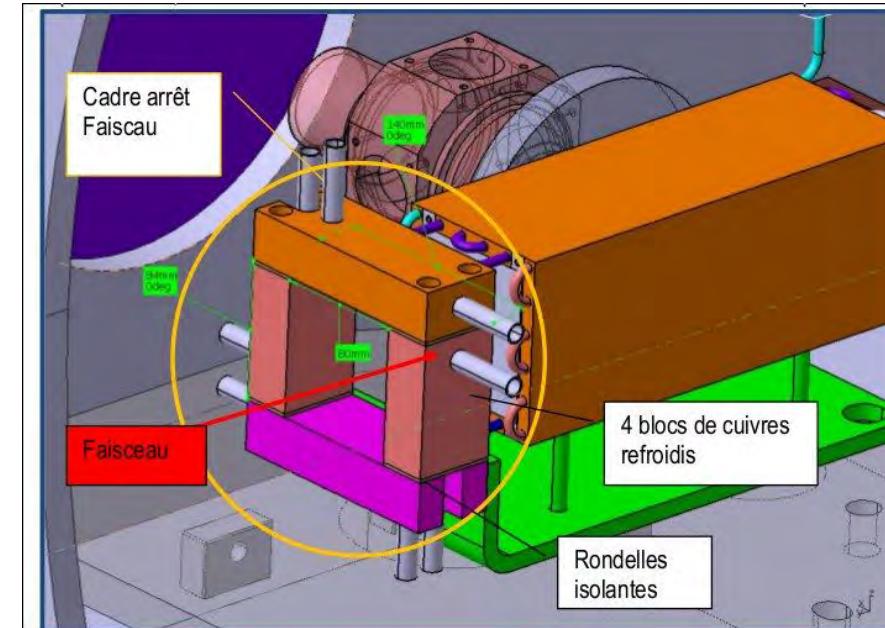
# Jaderná astrofyzika

responsible Jaromír Mrázek

Produkční terč ROBOT (Řež - Other Beams Other Targets)  
produkce izotopů  $^{14,15}\text{O}$ , které hrají roli např. při  
hot CNO cyklu a jejichž svazky jinak na SPIRAL2 nebudou.



SPIRAL2-CZ-OP investice







# Outlooks

